

Jet Propulsion Laboratory<br>California Institute of Technology

# **Low SWaP Radiation Sensor Development and Characterization at JPL**

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## **Topics**

- JPL's Radiation Monitors
- Sensors
- Summary

# **JPL's Radiation Monitor Heritage**



## **Environmental Sensors**



## **TID – RADFET**

16-bit 0V to 2.4V

- RADFET from REM was used
	- $-$  Size of chip carrier: 1.7 x 0.87 cm<sup>2</sup>, mass:  $< 1$  g

Example Usage

- Responsivity, Range, and Resolution
	- Responsivity: 0.2 mV/rad(Si) at low dose 0.02 mV/rad(Si) at high dose
	- Range: Sensor 1 Mrad (25 V)

ISS Environment < 250 krad (1 year)

Example Circuitry

- Resolution: dose resolution
	- (50 rad= 1 mV)

(dose during one orbit, 1.5 hours)

A/D converter: 16 bit



### **Calibration of TID sensor**



### **Calibration – built-in diode in TID sensor**



- TID 147 shows large deviation. The rest nine sensor show similar curves.
- The diode outputs show linear relationship with the temperature except for high temperature region (130  $\sim$  150 °C).
- TID sensors with similar temperature calibration curves (and similar I-V characteristics) were selected for the further tests.

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### **Low SWaP Adjustable Dosimeter**

- Developing with ASU (Hugh Barnaby and David Allee)
- A low size (<25 cm<sup>2</sup>), weight (<200 g), and power (<1 W) dosimeter system which is capable of adjusting its sensitivity (and maximum range) between 10 mrad (10 krad) and 1 rad (1 Mrad).
- Designing proto-type adjustable dosimeter system using REM Model TOT601B RADFETs combined with a dual slope Nyquist-rate A/D converter with built-in sync filter assembled with radiation hardened components.
- System will be able to sample between 2 rad and 1 Mrad of ionizing dose

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# **Internal Electrostatic Discharge Monitor (IESDM)**

The IESDM was designed to emulate a multi-layer circuit board. Very thin metallic layers (meshes) were inserted between the dielectric layers



## **Internal Electrostatic Discharge Monitor**

Example Monte Carlo (TIGER) simulation result – 1 MeV, 300 pA Show that insertion of thin metal layers does not change the charging profile



**Depth, mm**

## **Internal Electrostatic Discharge Monitor**





Dark conductivity: 3.1x10<sup>-16</sup> ohm<sup>-1</sup>m<sup>-1</sup> RIC coefficient: 2.0x10<sup>-16</sup> sec ohm<sup>-1</sup>m<sup>-1</sup>rad<sup>-1</sup>

#### Prototype IESDM



Test Setup



#### **SEU Sensor**

• Will use a commercial SRAM. This SRAM can be sensitized to SEUs by lowering  $V_{dd}$  (as was done for the RADMON). So we eliminate the  $V_{offset}$  adjustment channel, but keep the  $V_{dd}$  adjustment channel. If we only want two values (nominal  $V_{dd}$  value for read/write and lower sensitizing value for SEU "stare" time), the DAC can be replaced with a 2-1 AMUX feed by voltage dividers.



- RRELAX boxes flown on Clementine (1994)
- •Critical charge sensitivity to SEUs is tunable via Voffset and Vdd .

1.1 1.3 1.5 1.7 1.9 2.1 2.3 **Voffset (V)**

#### **SEU Sensor - RRELAX**

A signal level will strongly depend on shielding, integration time, pixel size, radiation hardness, threshold voltage as well as an exposed environment.

Refer to the previous experiment; RADMON on Clementine RRELAX (4 kbit)



Figure 6. Calibration curves and operational thresholds for the SRAM particle detector.

With the threshold voltage of 0.15 V, Protons with  $0.55$  MeV  $\lt E \lt 0.8$  MeV and heavier ions upset the RADMON.





Upset Rate of RADMON; Solar Flare: up to 6E4 SEU/30min Normal: ~ 2 SEU/30min

M. G. Buehler, G. A. Soli, B. R. Blaes, J. M. Ratliff, and H. B. Garrett, "Clementine RRELAX SRAM Particle Spectrometer," IEEE Transaction on Nuclear Science, vol. 41, pp. 2404, 1994.

#### **SEU Sensor**

Typical modern memory chip has 4 Mbit (1000 time higher than RADMON) and area of about 1 cm<sup>2</sup>.

Expected upsets calculating from RADMON data;

Solar Flare: up to 2E6 SEU/min

Normal: ~ 70 SEU/min

Calculation from Proton Environment:

Bin flux between  $0.5 \text{ MeV} < E < 1 \text{ MeV}$  (twice of that of RADMON case)

No Shielding

• Normal AP8MAX: 9.5E2 hits/min AP8MIN: 8.6E3 hits/min

• Solar flare From the left graph proton flux is  $\sim$  3E7/(m<sup>2</sup> s sr MeV) Converted to 5E5 /(cm<sup>2</sup> min)

Assuming 1% upset per measurement, 2E6 SEU/min can be done by a measurement per every second



Insoo Jun, "Ionizing Radiation Environment for the ISSA Low Temperature Microgravity Physics Facility (LTMPF)", JPL IOM 5132/5052-01-122, May 31, 2001

Figure 14. CREME96 High Energy Proton Differential Spectra for ISSA LTMPF **No Shielding** 

## **UV Sensor**

- A photodiode UVG12 from OptoDiode Corp. was selected
	- A direct measurement of the UV irradiance will be strongly interfered with by "contamination".
	- Measure broad band UV-VIS irradiance
		- Very small interference from contamination
		- Depends on the sun angle and shade
		- Measure Sun-Hour
	- $-$  Extract Ly-α irradiance from the measured sun hour and the reported daily Ly-α irradiance
		- The daily Ly-α irradiance is being measured in SOLSTICE instrument on SORCE satellite and reported in the website [http://lasp.colorado.edu/lisird/sorce/sorce\\_ssi/](http://lasp.colorado.edu/lisird/sorce/sorce_ssi/).
	- Some flight histories, Readily available
	- $-$  Spectral range: 140  $\sim$  1100 nm
	- $-$  Apply reverse bias (0 50 V) and read a current
	- Size: 1.0 x 1.0 cm<sup>2</sup>, mass: < 10 g
	- $-$  At normal, the expected irradiance is  $\sim$  0.1 W/cm<sup>2</sup>.

## **UV Sensor Calibration - Angle**



- Measured the output current at 1 AU irradiance
- Current measured as a function of the angle of the UV sensor.
	- Follows the cosine curve well
- Good for the local UV irradiance measurement

• Expected environment from SSP 30425

#### **TABLE 7.2-1 SOLAR ELECTROMAGNETIC RADIATION**



There are two possible cases depending on spectral range of interest.

1) 1 nm – 10 nm (1.2 keV – 0.12 keV) Difference with and without flare is only few times. Too many hits to use a pulse counting mode.

#### → Current Measurement

2) 0.1 nm – 1 nm (12 keV – 1.2 keV)

Difference with and without flare is several orders of magnitude. Current due to electron and proton is larger than the current due to xray in normal condition (without flare)

 $\rightarrow$  Pulse Counting

#### • **Contamination from electrons and protons.**

(From AE8MAX and AP8MIN and stopping powers from NIST) Assuming the detector (Si) thickness of 100 µm

Energy deposition due to electrons: 6.6 E6 keV/(cm<sup>2</sup> s) Energy deposition due to protons: 2.8 E5 keV/(cm<sup>2</sup> s) Total  $\sim$  6.9 E6 keV/(cm<sup>2</sup> s) which is  $\sim$  1 E-5 W/m<sup>2</sup>

Case 1), current measurement mode,

the contamination due to the electrons and protons will be less than 20 %. Case 2), pulse counting mode,

the contamination from electron and proton can be distinguished by different pulse height.

X-ray: less than 10 keV Electron: 50 – 100 keV Proton: 100 – 1000 keV



http://hesperia.gsfc.nasa.gov/hessi/flares.htm

Case 1) Sensor: AXUV100 from OptoDiode Inc, sensor area 1 cm<sup>2</sup>, Responsivity 0.28 A/W Expected current: 1.4 nA

Case 2) Sensor: Si-PIN from Amptek, sensor area 6 mm<sup>2</sup>, thickness 500  $\mu$ m, 1.0 mil Be Window Maximum pulse counting rate: 1 E6 pulses/sec Expected pulse rate at solar flare: 3E5 pulses/sec



#### **Camera**

- Optional module
- Low mass, volume, and power
- Could include Gumstix (image processing) and local memory within the module
- Candidate detectors:
	- 5MP Aptina MT9P031
	- Mouser 931-LI-5M05CMAF
	- 0.3 MP Aptina MT9v032 (MarCO's detector)
- Autofocus vs fixed lens



*based camera module*

Lexar

*MarCO camera (Gumstix module and 0.3 MP Aptina-based camera module)*



## **Summary**

- At JPL, compact radiation monitor activities have been going on.
- Several sensor were considered and calibrated.
- In a future, we hope that every JPL missions have the radiation monitor.

# **THANK YOU!**

# **QUESTIONS?**



### **Test Results – Temperature sensor**





- All ten sensors show very similar calibration curves.
- There is a small discontinuity at 20 ˚C.
	- Most likely due to configuration difference between high and low temperature calibration setups.
	- Final temperature calibration is needed as the final flight configuration: Test As You Fly (TAYF)

## **UV Sensor – UVG12**

- Responsivity, Range, and Resolution
	- Average Responsivity: 0.25 A/W
	- Expected maximum signal (when the surface is normal to Sun): 3 mA
	- Range: Sensor 10 mA (linear up to 3.1 mA) Environment 3 mA

