

Europa Ice Penetrator

Towards a Hybrid Ice-Penetrating System

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*Jason C. Swenson and †Alexandre M. Kling

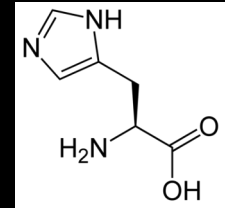
* USRA Education Associate, Mission Design Division (Code RD), NASA Ames Research Center Bldg N202 Room 200.

† USRA Education Associate, Space Science and Astrobiology Division (Code SST), NASA Ames Research Center Bldg N245 Room 280.

Accessing the Subsurface Constituents on Europa

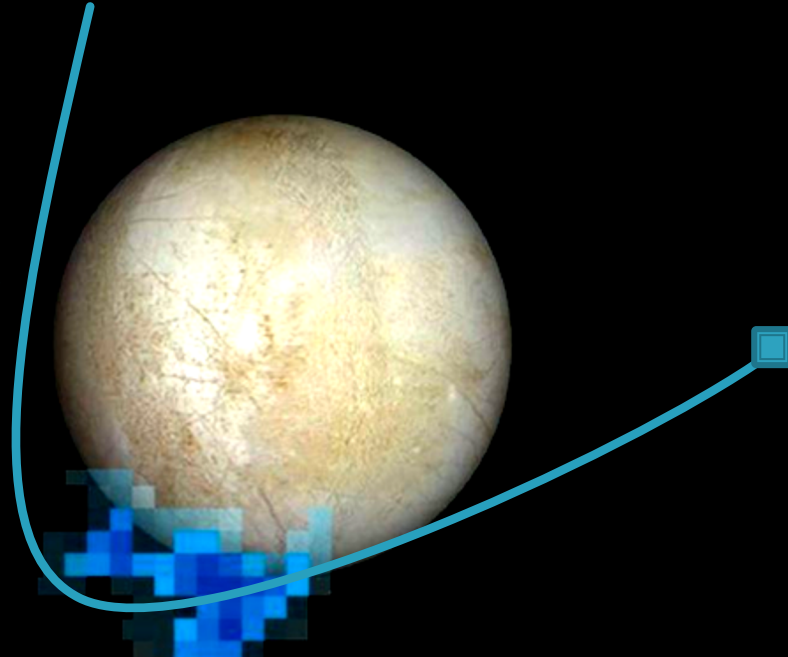
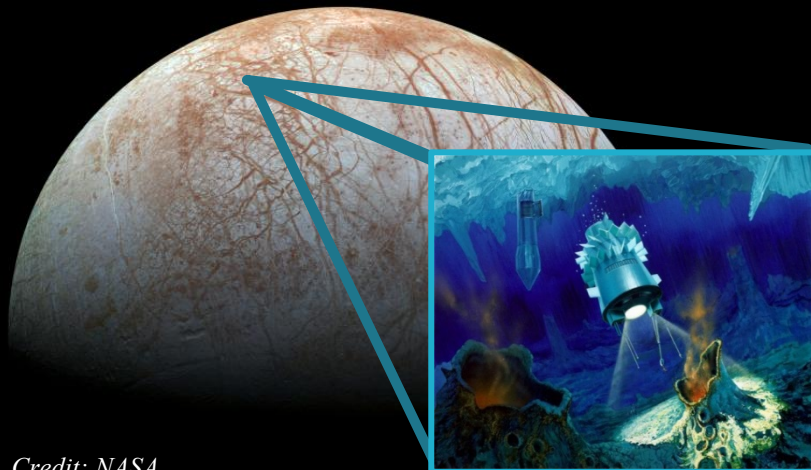
A Hypothesized Subsurface Ocean

Evidence for life?



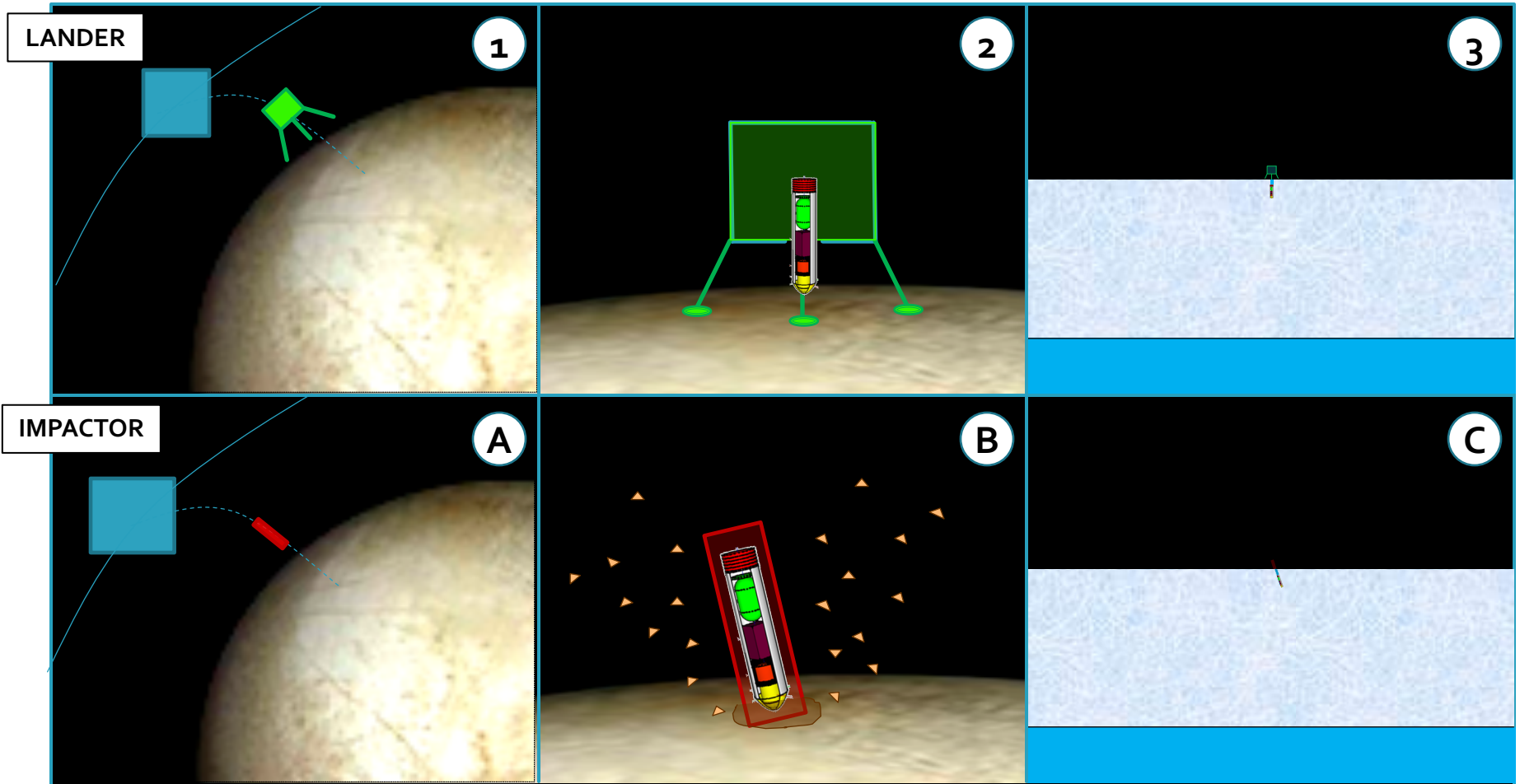
Method 1: Penetrate the surface

Method 2: Sample and analyze the plume



Credit: NASA

ConOps Possibilities



** For conceptual purposes only. Not all potential mission sequences are discussed.*

Use of Small Spacecraft?

Earth-based techniques

- Non-applicable due mass and volume constraints



Courtesy Y. Bar-Cohen, K. Zacny

Flagship-class mission

- Is this the way to go? (Perhaps...)



"Aliens of the Deep ", James Cameron

Small(er) spacecraft

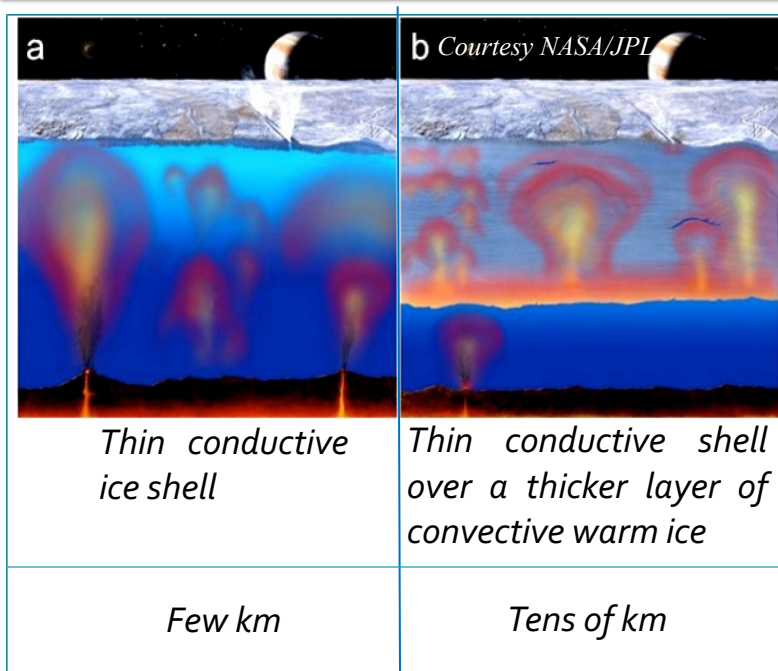
- Can we do something? (requires understanding of payloads involved)



ISSC 2015

Low-cost scientific studies can be enabled with strict constraints on spacecraft mass and volume.

Global Ice Shell Discrepancies

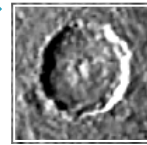


Methods



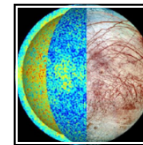
Galileo's Doppler & Magnetometer Data

- Internal structure, thickness of H₂O layer



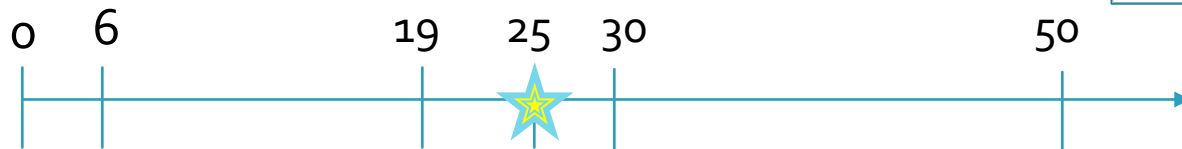
Observation of geological features

- Domes, pits, flexure, and craters



Thermal modeling

- Tidal heating, heat dissipation



6-30 km: supported topography

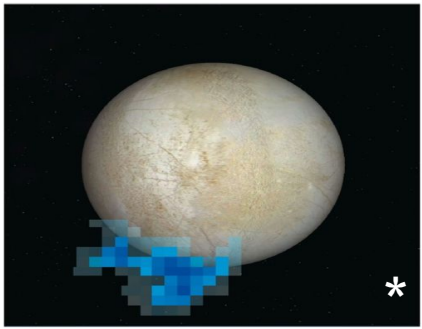
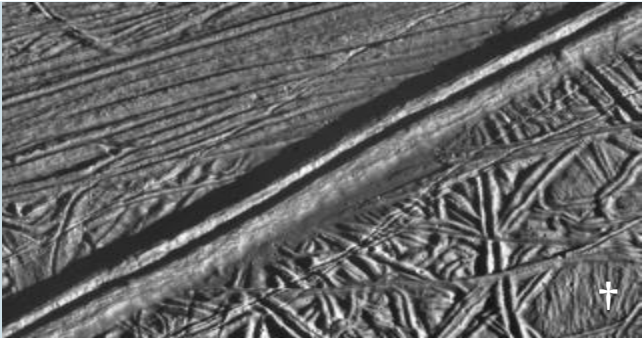
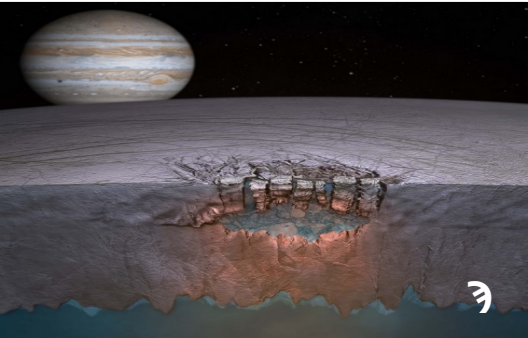
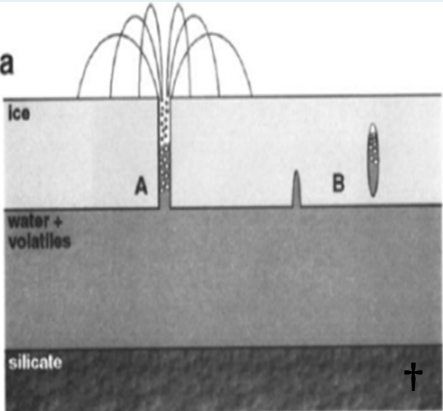
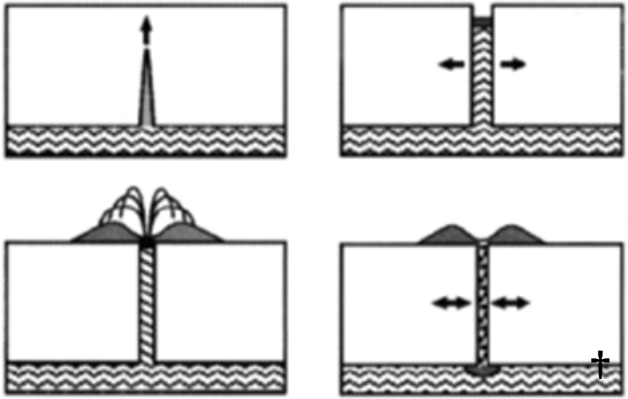
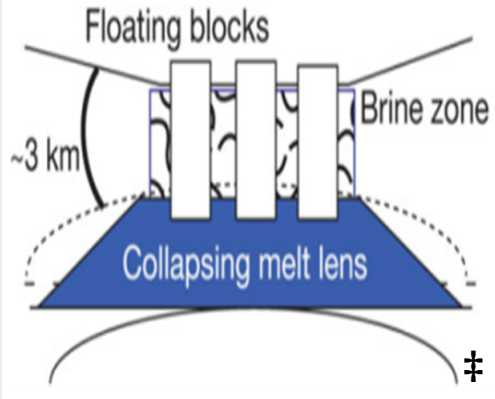
25-50 km: thermal modeling

19 - 25 km: impact craters

Estimation

25 km is a *plausible estimate*

Local Ice Shell Discrepancies

	Plumes	Ridges	Chaos Terrain
Observations			
Models			

* Courtesy of NASA, ESA, and L. Roth

‡ Courtesy of University of Texas, Austin

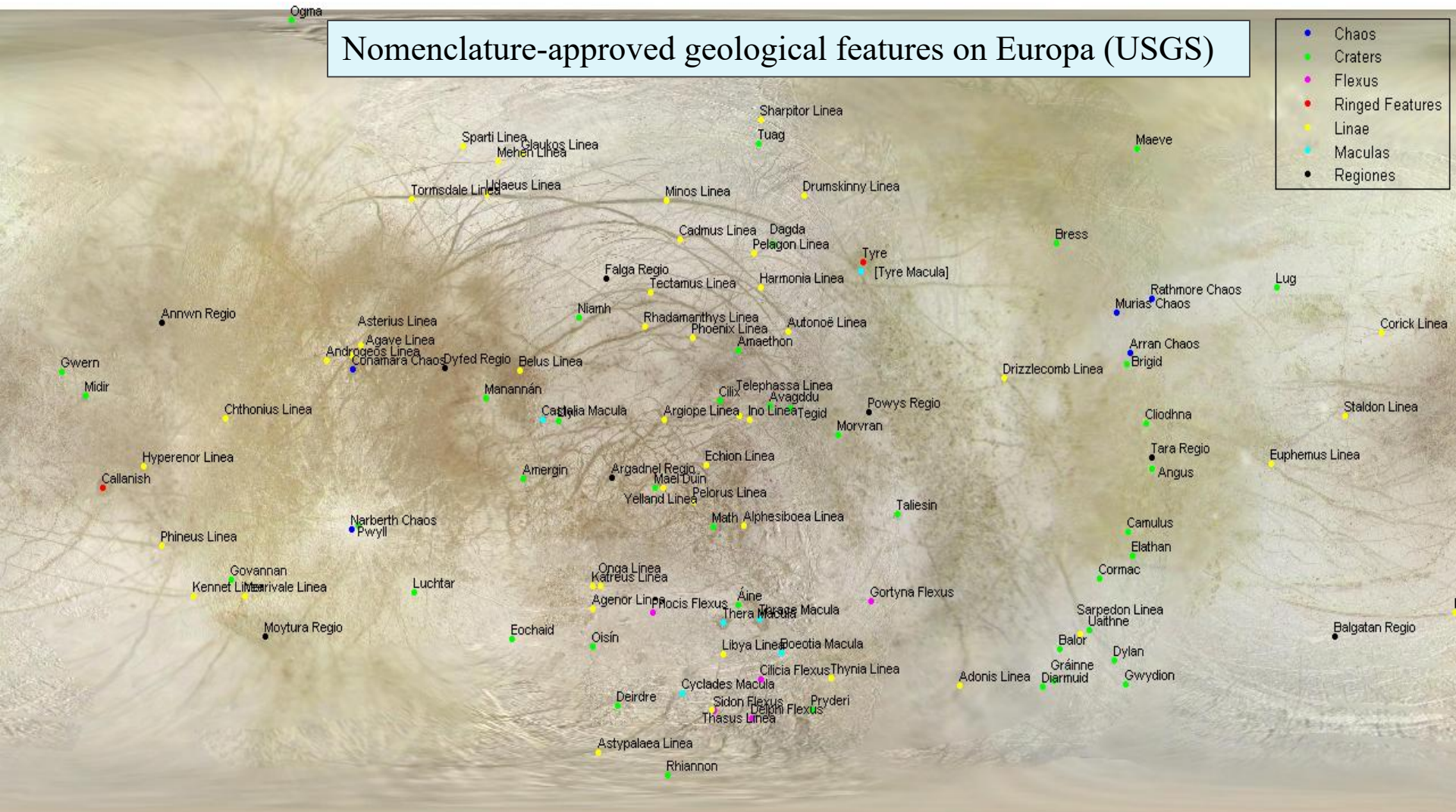
† R.T.Pappalardo, 1999. Does Europa Have a Subsurface Ocean? Evaluation of the Geological Evidence. *Journal of Geo. Research*, Vol 104, pp. 24.015-14;055.

‡ Schmidt, B. E et al., 2011. Active Formation of 'Chaos Terrain' over Shallow Subsurface Water on Europa. *Nature*, 479(7374), pp. 502-504.

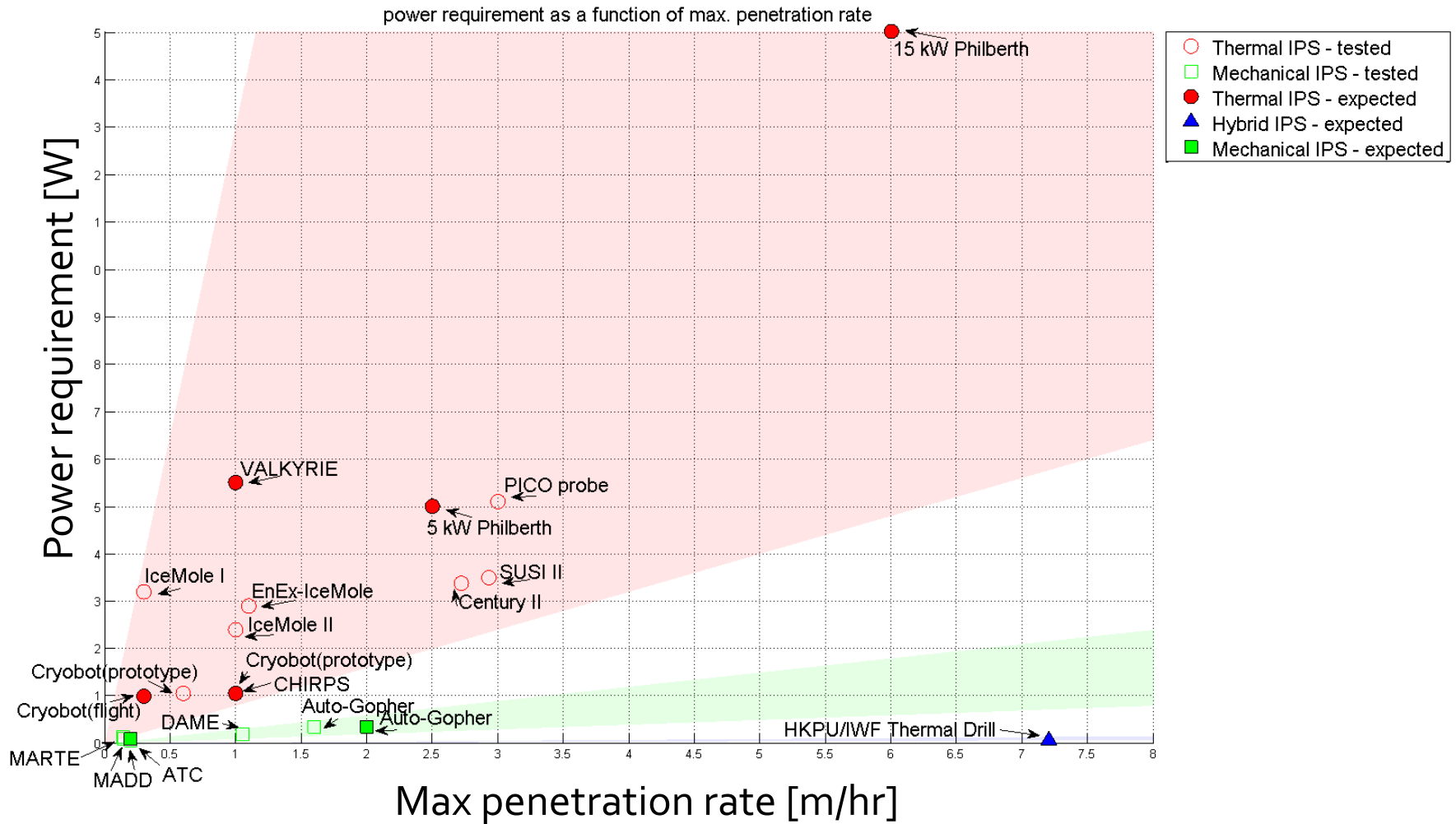
Deployment Sites

Nomenclature-approved geological features on Europa (USGS)

- Chaos
- Craters
- Flexus
- Ringed Features
- Lineae
- Maculas
- Regiones



Ice-Penetrating System (IPS) Technology Review



Dimensional Analysis

Presence of Salts

Presence of Rocks

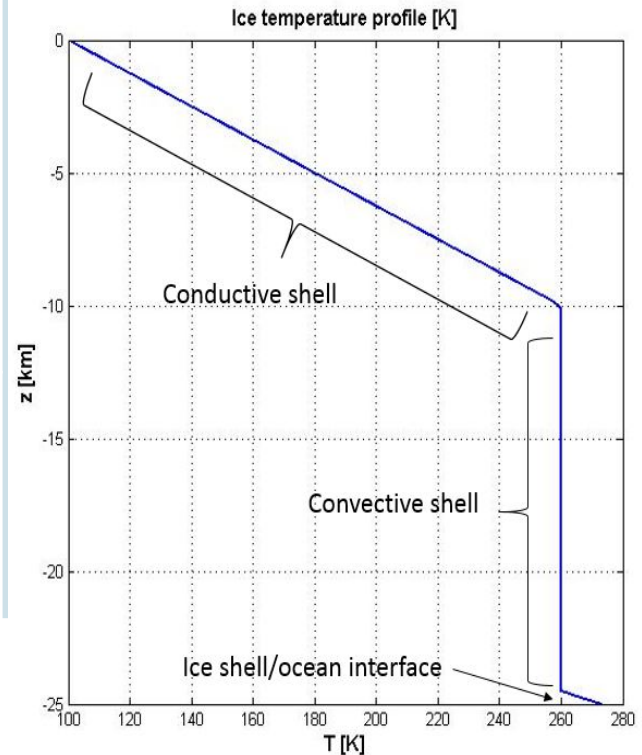
Temperature Profile

Hypothesis

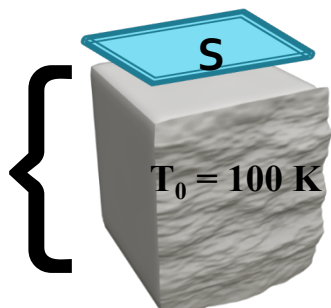
- ~85% of Europa's surface is ice or frost
- Probable non-ice candidates: (MgSO₄·6H₂O, MgSO₄·7H₂O,) or Na₂CO₃·10H₂O
- Concentrations: 1%, 5%, 20%??



- Presence of rocks unlikely
- Ecliptic impact probability of Europa relative to Jupiter: 6.6×10^{-5}

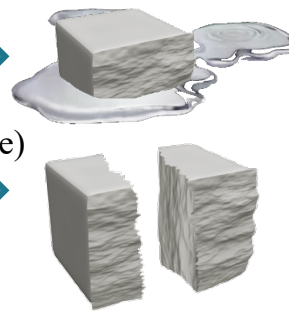


L meters



Melting energy:
 $E_{\text{melt}} \approx 450 \text{ MJ/m}^3$ (surface)

Specific cutting energy:
 $E_{\text{cut}} \approx 5 \text{ MJ/m}^3$



- Energy, E , required to melt L meters of ice
- Power \rightarrow time required to deliver E
- penetration rate

Thermal IPS

Cryobot: Time to penetrate 25km, $\eta=0.7$

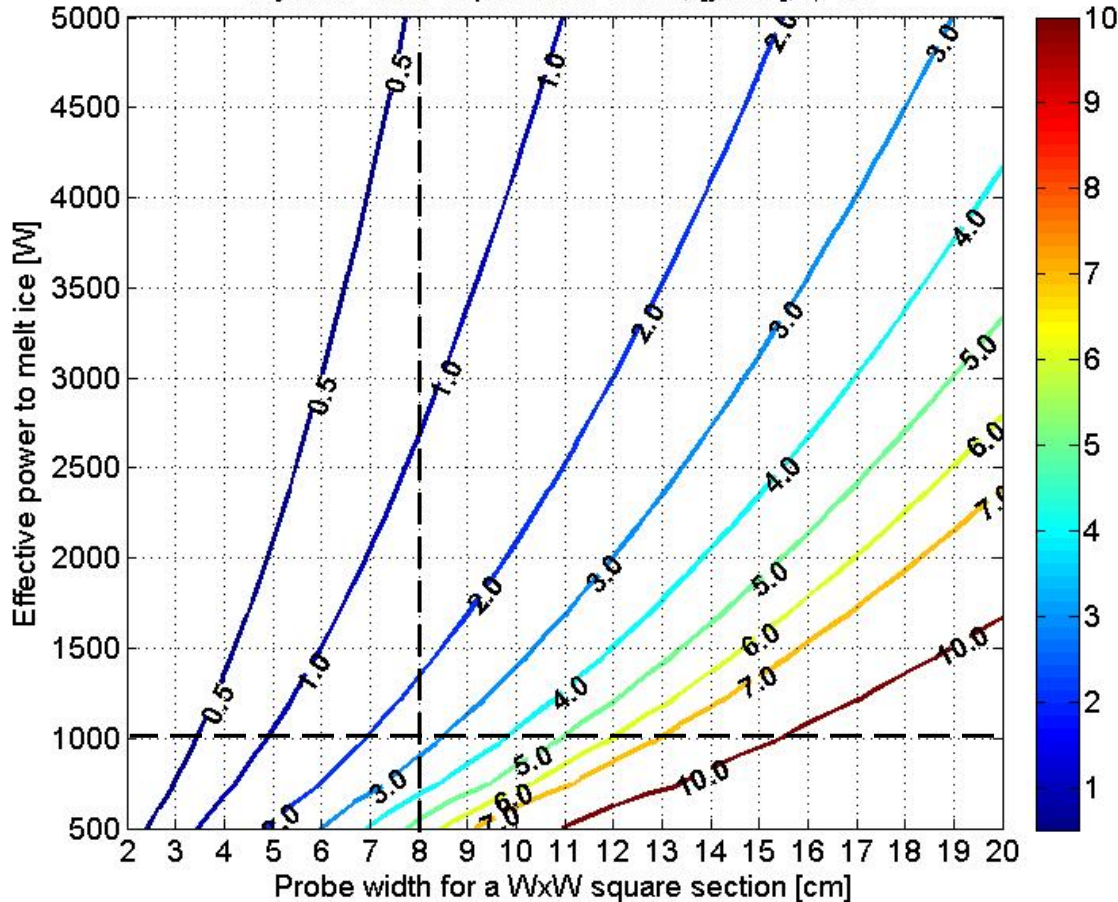
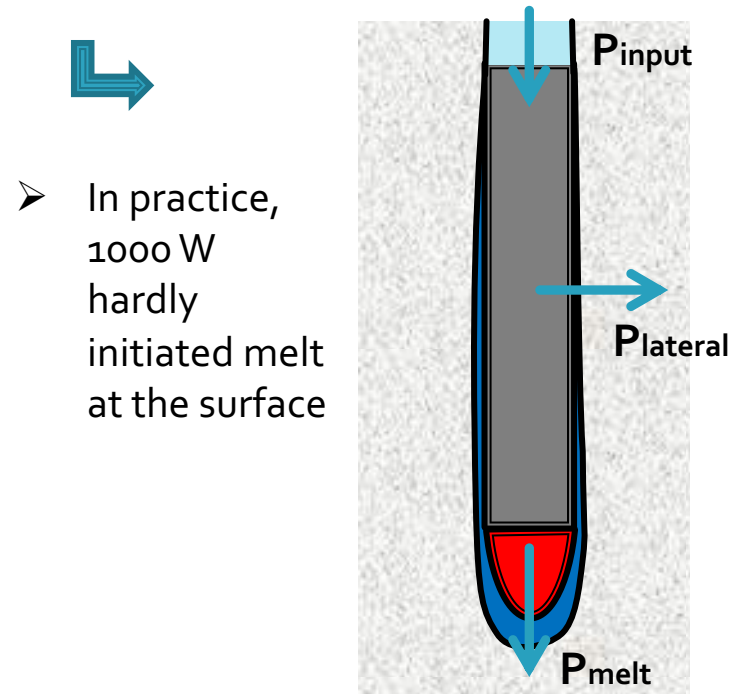


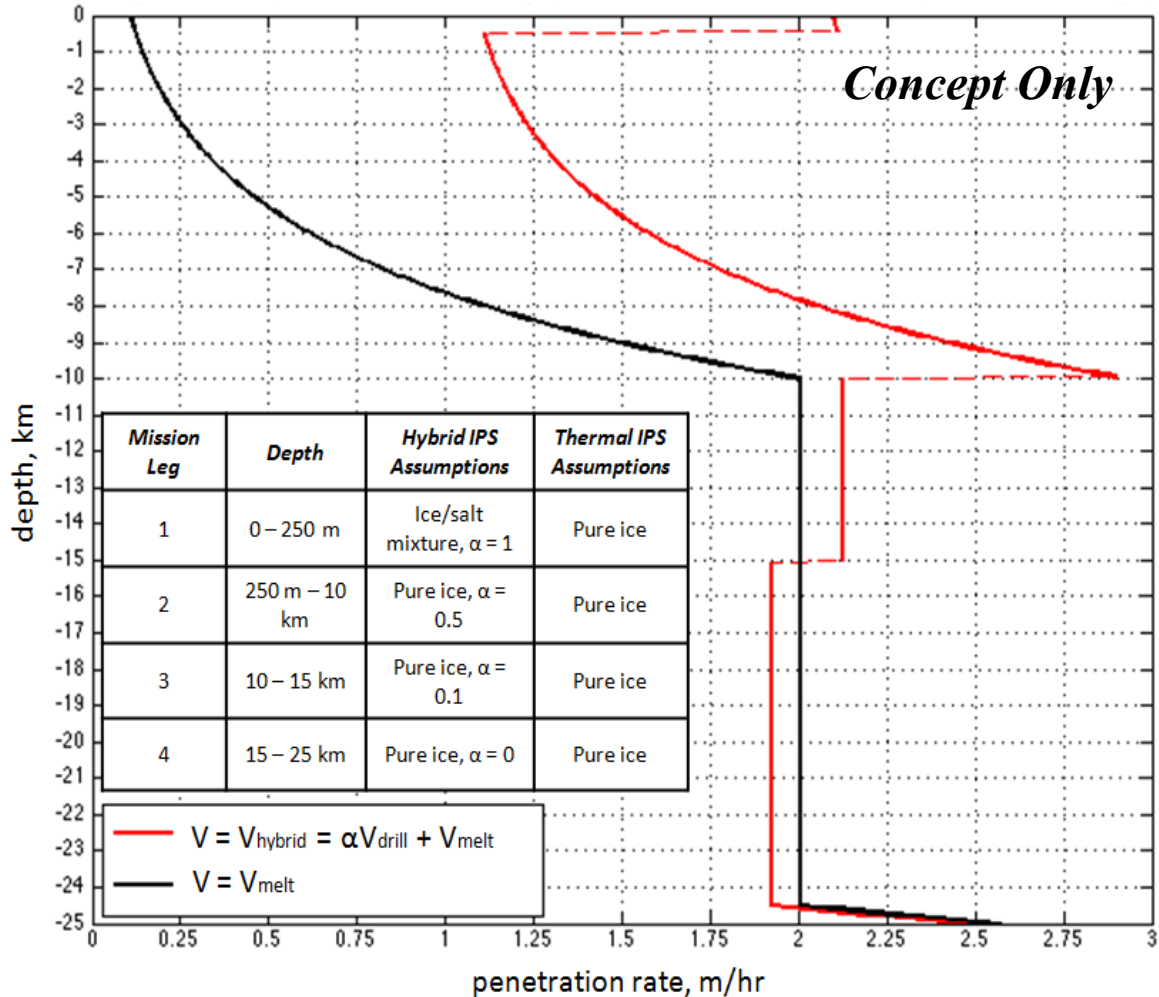
Chart derived from equations in the following reference: Aamot, H. W. C., 1967. *Heat Transfer and Performance of a Thermal Probe for Glaciers*; Thermal IPS chart does not consider lateral heating.

Lateral Heating Considered?	Input Power	Total Penetration Time (TPT)
NO	> 1 kW	< 2.75 years
YES	2 kW	~4.5 years

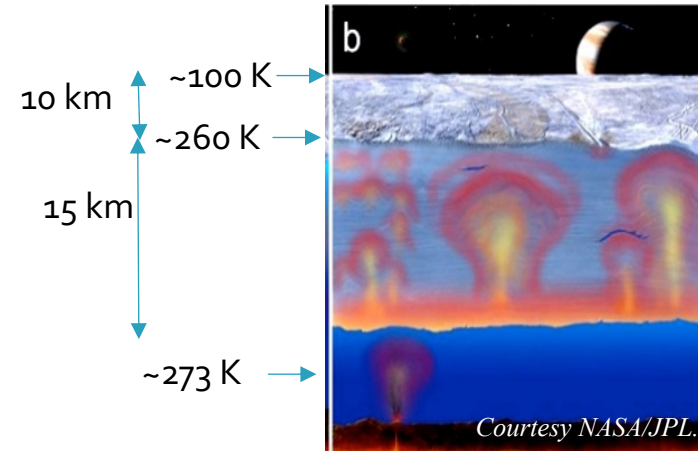


Hybrid IPS

8 x 8 cm Hybrid IPS (70 – 2000 W, 1.60 yr) vs. 8 x 8 cm Thermal IPS (2000 W, 4.36 yr), 25 km



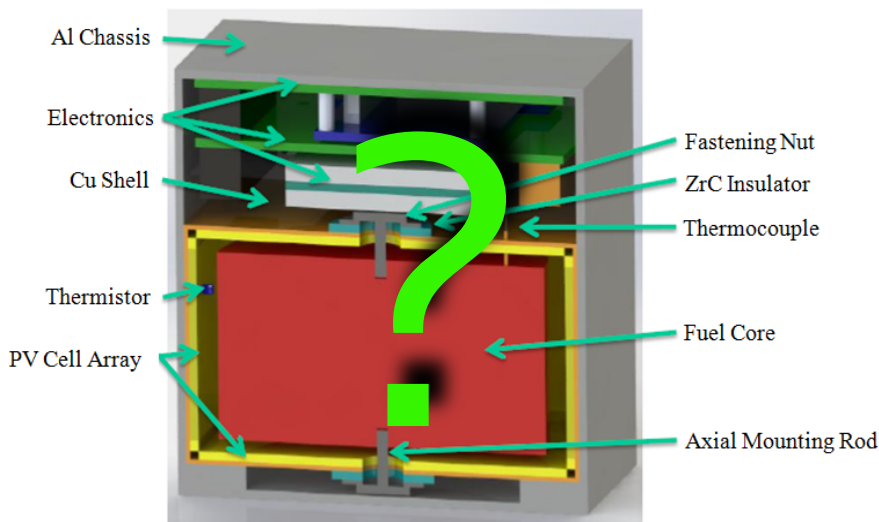
- Melting at the surface is highly inefficient
- Mechanical drilling allows faster penetration rate at the surface
- If convection is initiated within the crust, the ice shell is thick and temperatures rise quickly
- (In our model, $dT/dz = 16 \text{ K/km}$)



System Considerations: Power Distribution and Communication

How to **power** an IPS traversing 25 km?

- Surface-based power
 - Large mass requirement
 - Tethers limit achievable depth
- Internal power
 - RPS (e.g. RTPV) could provide 1.4 kW for a melthead / drill actuator

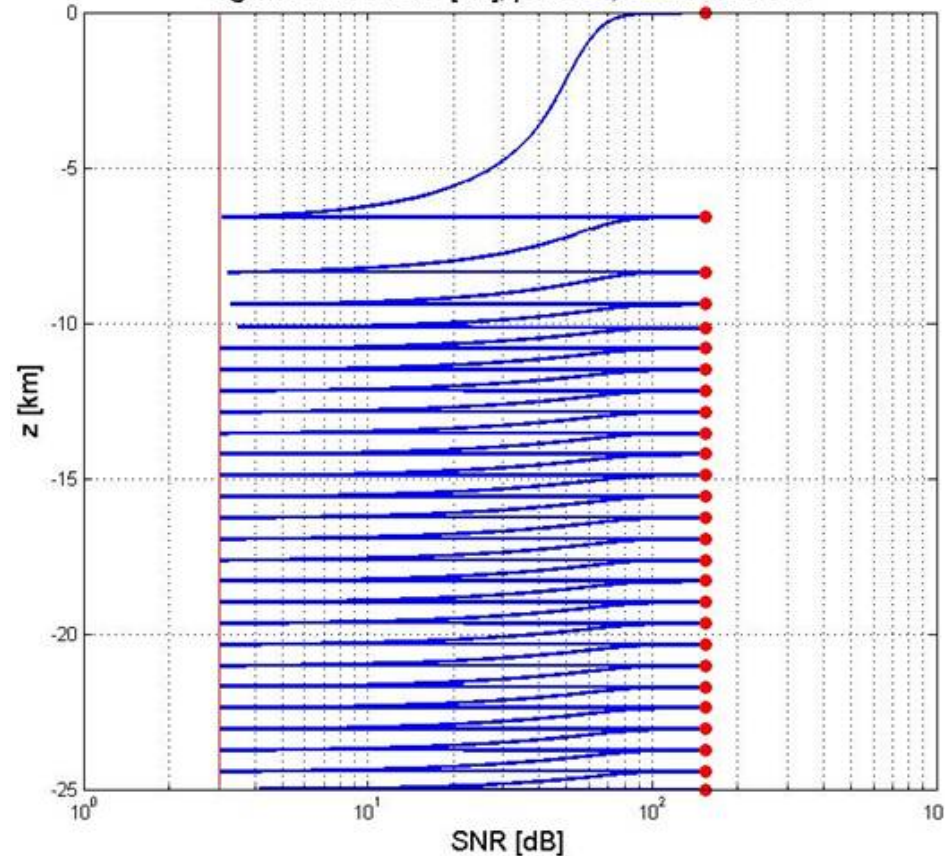


A. Goel, K. Schillo, B. Franz and S. Reddy, 2014. "Radioisotope Thermophotovoltaics (RTPV) Flight Demonstration," Idaho Falls, ID, 2014.

How to **communicate** through 25 km of ice?

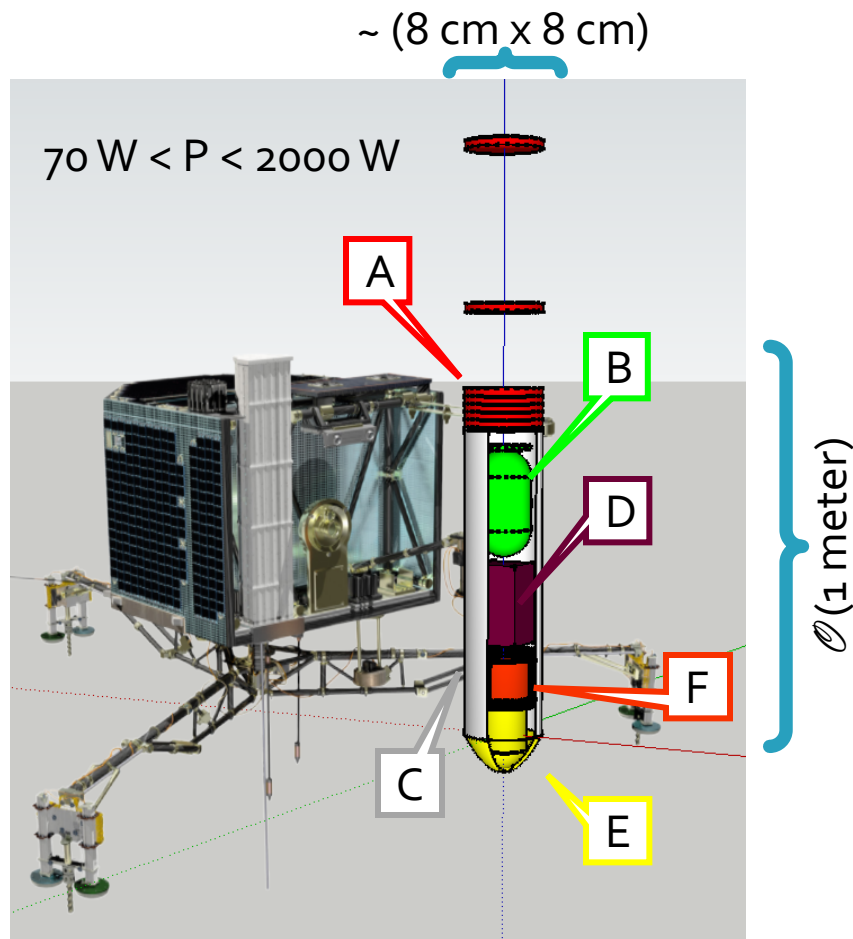
➤ Mini RF Ice Transceivers

Signal to noise ratio [dB], pure ice, 27 transceivers



Bryant, S., 2002. *Ice-Embedded Transceivers for Europa Cryobot Communications*. Big Sky, Montana.

Concept: Hybrid IPS for 25 km Ice Shell

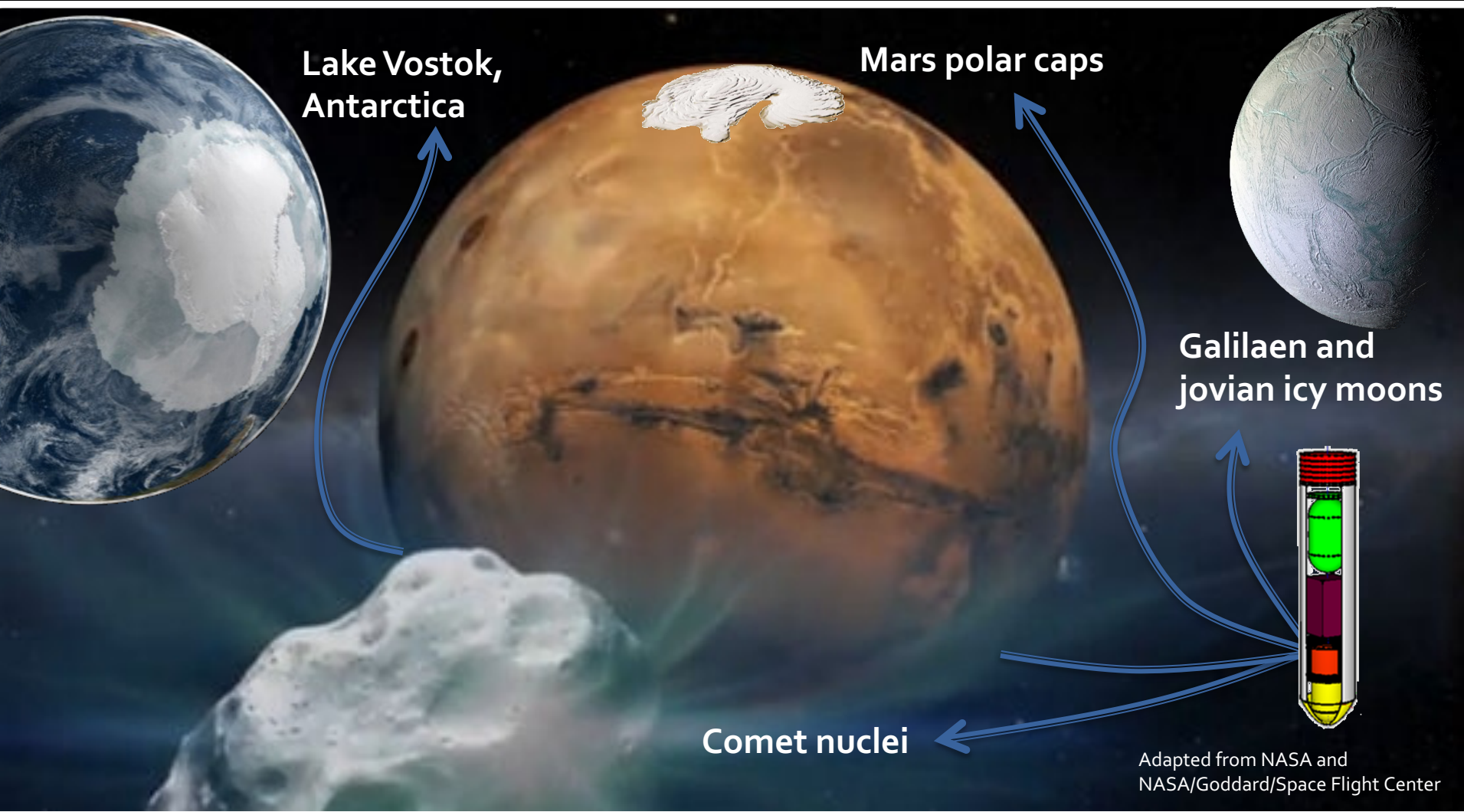


	Component	Remarks
A	Mini RF Ice Transceivers	Robust to ice shell motion and borehole refreezing
B	Science Payload	- "Lab on chip"(€) - Conductivity, Temp. and Depth Sensor (δ) - underwater vehicle
C	IPS Body	Lateral heating
D	Control Unit & Sensors	- IMU sensor - Ultrasound Sensor - Thermal/IR Imager
E	Melthead / Drilling Actuator	- Rotary drill - Hammer drill
F	Radioisotope Power Source (RPS)	If penetration depth >3km, otherwise tether may be possible

€ Butterworth, A.L. et al. Lab-on-a-Chip Organic Analyzer: Instrumentation and Methods for Detecting Trace Organic Molecules and Amino Acid Chirality in Planetary Science. 46th LPSC, 2015.

δ Jonsson, Jet al. Chip-Based Salinity Measurements for Small Submersibles and Biologgers. International Journal of Oceanography, 2013.

Need for an IPS Able to Penetrate Ice-rock Mixtures



Conclusion

- Hybrid (melting and drilling) system would enable drilling through cold surface ice and salts
- A 70 – 2000 W hybrid IPS unit can theoretically complete its mission in approximately 1.6 years, nearly three times faster than a 2000 W thermal IPS with lateral losses included and of the same cross sectional area
- RPS technology is required to achieve large ($\gg 3$ km) penetration depths; 27 ice-embedded mini RF transceivers are required for communications at 25 km
- Potential for a hybrid IPS to explore other icy bodies (moons, comet nuclei, Mars' polar caps, etc.) or study subglacial environments (Antarctica, Arctic, etc.)

