National Aeronautics and Space Administration



The Plume Chaser Mission:Two-Spacecraft Search for Organics on the Dwarf Planet Ceres

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Ceres

- Aphelion: 2.9773 AU
- Perihelion: 2.5577 AU •
- Orbital period: 4.60 yr ٠
- Avg. orbital speed: 17.905 km/s ٠
- Mean radius: 476.2 ± 1.7 km ٠
- Mean density: 2.077±0.036 g/cm3 ٠
- Surface gravity: 0.28 m/s2
- Escape velocity: 0.51 km/s ٠
- Mean surf. Temp.: 168 K ٠
- Max surf. Temp.: 235 K ٠













Why is Ceres so interesting?

- Differentiated structure with a anhydrous silicate core with water-rich mantle
- The relatively rare CM chondrites are thought to originate from here
- Spared from the effects of the Late Heavy Bombardment (LHB)
- HST detected up to 8 features releasing 10^26 molecules (~6 kg) of material / s
- HSO found water vapor in the vicinity of 2 mid-latitude dark regions
- Thermal models suggest liquid water below the surface
- Could there be cryovolcanism and liquid water bodies under the frozen crust?





Why another mission to Ceres?

- Dawn arrives close to aphelion, while plume activity was observed at perihelion
- Dawn instruments are not capable of measuring needed critical isotopic composition of volatiles
- Dawn has no impactor to enable
 - the release subsurface volatiles
 - imaging before and after an impact event



→ Plume Chaser has a different, but complementary mission concept to Dawn







Science Objectives

Primary Science Objectives

- A.1 Search for evidence of organics in the Ceres water vapor/plume
- A.2 Determine the column water content and location of highest density
- A.3 Determine the elemental composition (C, H, O, N) of any carbon rich plume particles
- A.4 Characterize the particles in the plume for dissolved salts

Secondary Science Objectives

- B.1 Characterize the silicate dust in the plume and relate it to other solar system objects, such as meteorites and comets
- B.2 Characterize the mechanical, chemical and geological properties of the crust by observing the impact site before and after impact
- B.3 Look for surface compositional changes on the surface associated with venting; characterize composition differences involving water or possible organics





Instrumentation

- ENIJA Mass Spectrometer TRL 4 Particle size: 10 nm – 100 um Particle mass: 1 – 2000 amu
- NASA ARC/Draper Lab. NIR spectrometer

TRL 4

Range: 1.6 – 3.4 um

Resolution: 0.15 um

GomSpace NanoCam C1U camera

TRL 9 Spectrum: 400 – 1000 nm FOV: 9.22 deg

Resolution: < 80 m/pixel from 650 km



University of Stuttgart



NASA Ames

GomSpace



Mission Concept: Sniff, Wait, and Impact

- Sniff
 - Venting from the surface of Ceres is common
 - Permanent exosphere
 - Densest parts will be localized and studied
- Wait
 - Venting is not common on Ceres, but likely at perihelion
 - S/C must remain in orbit until a plume event occurs and can subsequently be maneuvered through it
- Impact
 - Venting is rare on Ceres
 - Sporadic occurrences under extremely favorable conditions
 - Trace exosphere might be beyond detection limits
 - To guarantee scientific return, impactor will create a plume to fly through







The Two Spacec

- Same design for both S/C Size: 40x40x30 cm^3 Dry mass: 40 kg Wet mass: <190 kg Fit within ESPA Grand frame
- 3D printed Ti tank/structure
- MMA Design solar array
 200 W EOL
- Propulsion

CubeSat Ambipolar Thruster (CAT) ISP: 1200 Size: 1U Propellant: Iodine



University of Michigan





Trajectory Design

- Starting from GTO
- Plume Chaser
 - Dry mass: 43 kg
 - Propellant: 120 kg
 - Overall duration: 4.8 yrs
 - Delta-V: 14.4 km/s
 - 1 year science orbit before decommision

Discoverv

Innovations

Solution

- Plume Maker
 - Impact mass: 41 kg
 - Propellant: 99 kg
 - Overall duration: 2.1 yrs
 - Delta-V: 14.5 km/s
 - Impact speed: 9.3 km/s





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Impacting

- To create a crater
 - Speed: 9.3 km/s
 - Mass: 40.9 kg
 - Energy: 1764.5 MJ
 - Crater
 - > 10 m wide
 - > 3 m deep
- To create a plume
 - Reach above 200 km
 - With particle velocity 750-2000 m/s



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Conclusions

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