Ice Cube Lunar Orbiter with BIRCHES (Broadband InfraRed Compact High-Resolution Exploration Spectrometer)

NEXTSTEP LunarCubes Mission and Instrument Concept

P.E. Clark, IACS/CUA/NASA GSFC and Flexure Engineering, Science PI

B. Malphrus, Morehead State University, PI

Morehead State University: B. Twiggs, Jeff Kruth, Kevin Brown, R. McNeill, B. Kroll
NASA/GSFC: A. Mandell, R. MacDowall, C. Brambora, D. Patel, S. Banks, D. Folta, P. Calhoun, P. Coulter
Busek: K. Hohman, V. Hruby

Next Step Selectee Announced March 30, 2015!

Why Lunarcubes?

Using the Cubesat paradigm to build user requirements driven ‘pathfinders’ for low-cost multi-platform mission concepts that will ultimately provide next generation exploration through the use of temporal and spatially distributed measurements.

Providing access to deep space via the Moon as nearby analogue, technology testbed, and gateway to the solar system.

Providing a low-cost alternative for high science yield missions at a time of declining funding and increasing costs for conventional missions.

Taking advantage of the decade long evolution of the cubesat model from standardized kits to science-driven, multi-institutional, multi-platform collaborations for LEO applications.

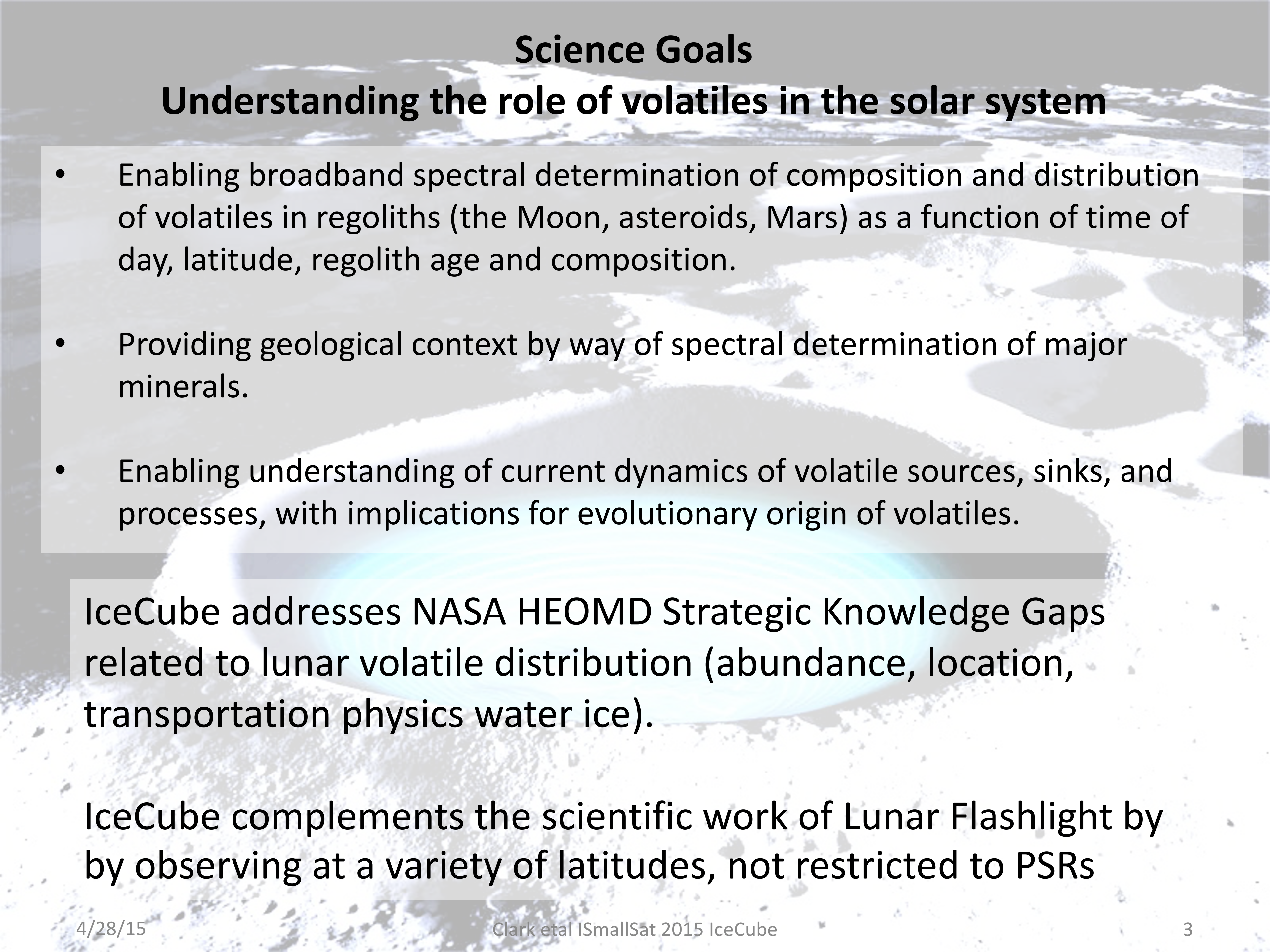
Examining the use of cubesat hardware/software for missions that are a representative cross-section of lunar, Mars, and other applications at varying degrees of difficulty (flyby, probe, orbiter, lander).

identifying modifications and new technology needed to support a science-driven deep space mode.

Looking for NASA to expand the CubeSat Launch Initiative which provides launch opportunities for cubesats to LEO as secondaries at no cost, to GEO and beyond.

designing a deep space prototype bus, and prototype for a lunar orbiter missions.

Building on the exploding interest in cubesat as seen in growing popularity of our LunarCubes Workshops over the last 3 years.



Science Goals

Understanding the role of volatiles in the solar system

- Enabling broadband spectral determination of composition and distribution of volatiles in regoliths (the Moon, asteroids, Mars) as a function of time of day, latitude, regolith age and composition.
- Providing geological context by way of spectral determination of major minerals.
- Enabling understanding of current dynamics of volatile sources, sinks, and processes, with implications for evolutionary origin of volatiles.

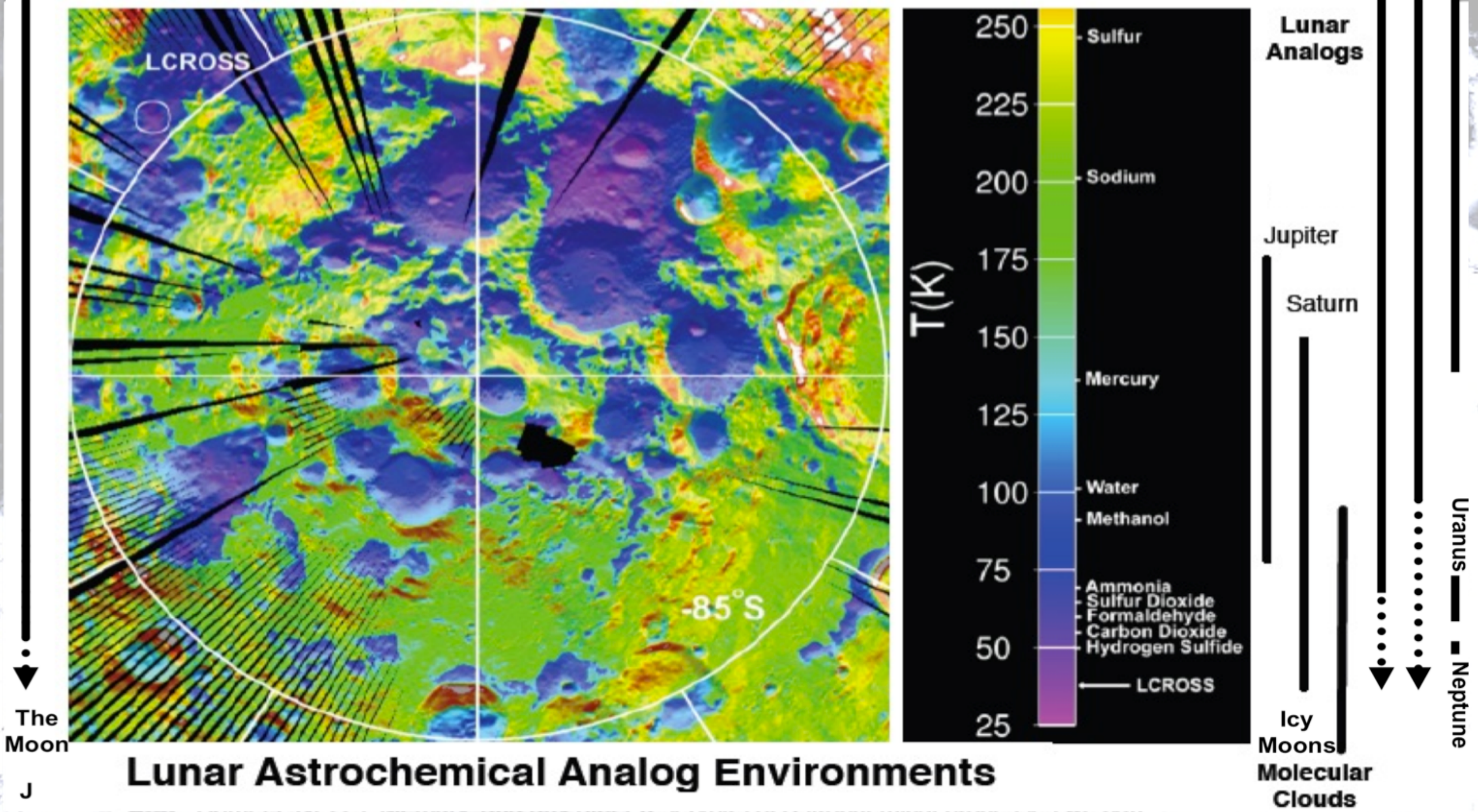
IceCube addresses NASA HEOMD Strategic Knowledge Gaps related to lunar volatile distribution (abundance, location, transportation physics water ice).

IceCube complements the scientific work of Lunar Flashlight by by observing at a variety of latitudes, not restricted to PSRs

The Extreme Lunar Environme

Thermal Extremes
Unmitigated Space Radiation
Abrasive Dust

Location	Day Temperature and Length	Night Temperature and Length
Low Latitude	400K, 14 days	120K, 14 days
Near Polar	220K, permanent	<25K, permanent



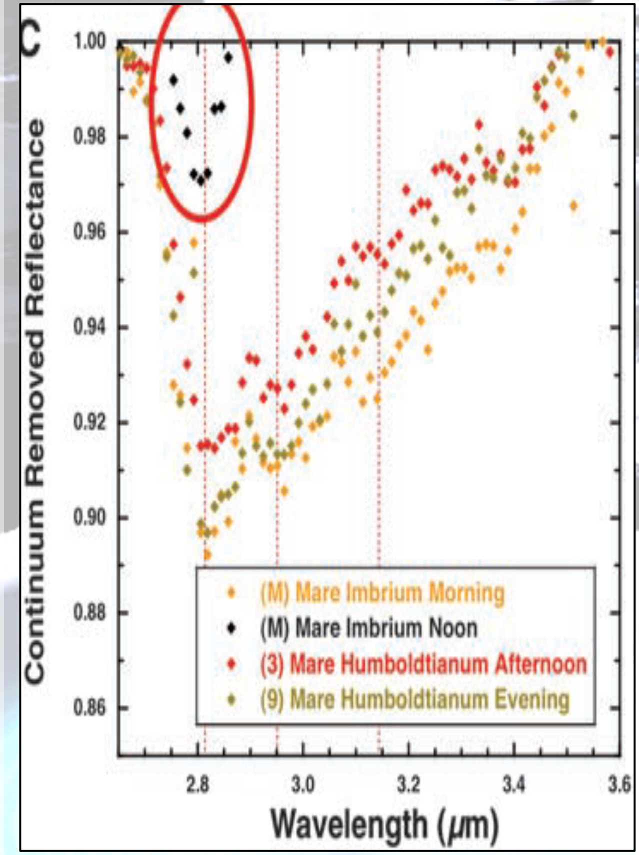
Lunar Astrochemical Analog Environments

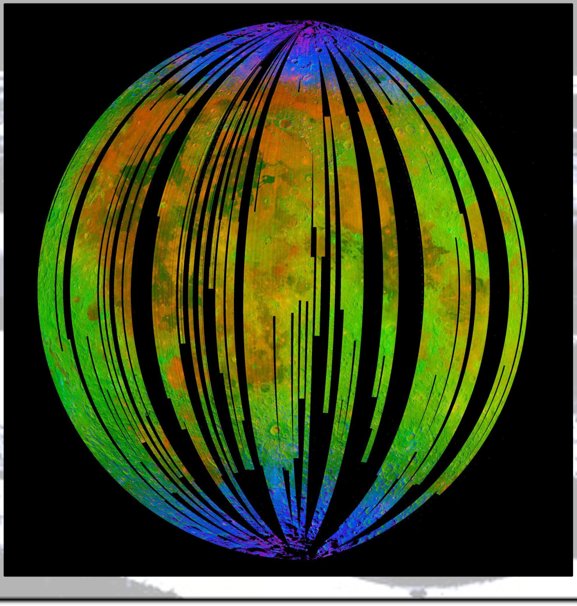
While M3 provided a 'snapshot' mosaic of lunar nearside indicating surface coating of OH/H_2O (blue) near the poles,

Early evidence for diurnal variation trend in OH absorption (Sunshine et al. 2009)

LCROSS provided evidence of additional subsurface volatiles.

IceCube will extend 'snapshots' to geospatially linked time of day and latitude coverage.





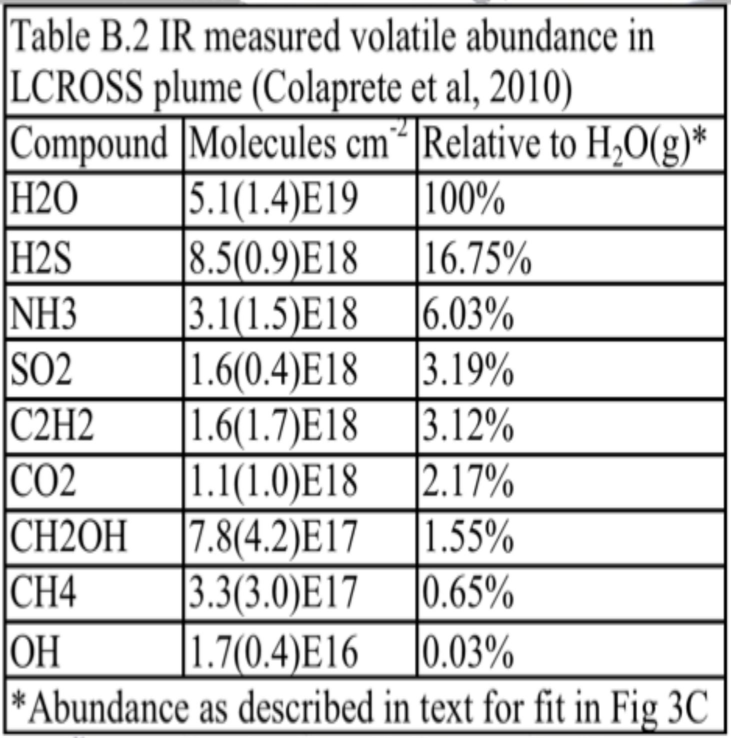
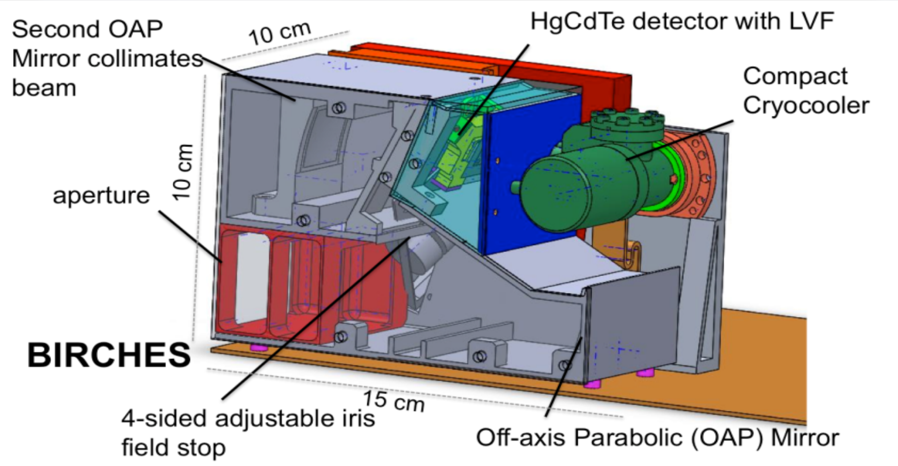


Table B.2 IR measured volatile abundance in LCROSS plume (Colaprete et al, 2010)

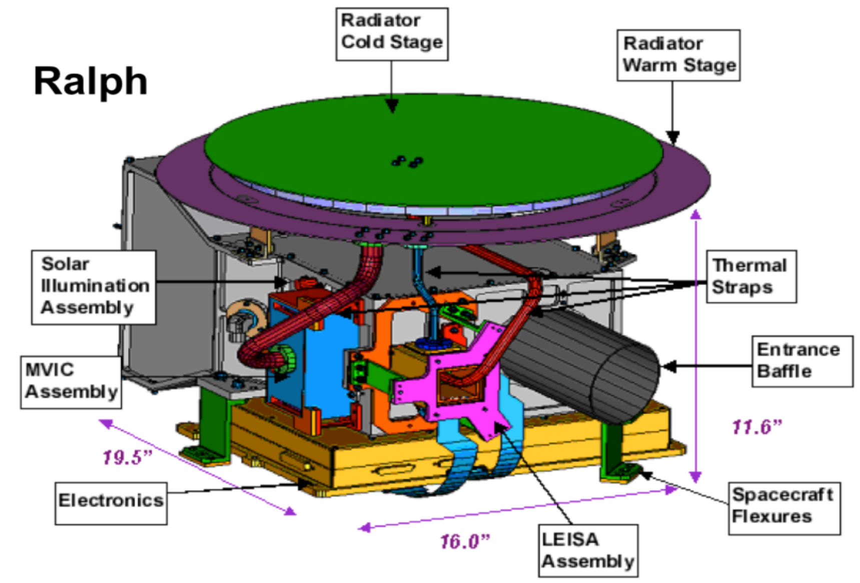
Compound	Molecules cm^{-2}	Relative to H_2O(g)*
H2O	5.1(1.4)E19	100%
H2S	8.5(0.9)E18	16.75%
NH3	3.1(1.5)E18	6.03%
SO2	1.6(0.4)E18	3.19%
C2H2	1.6(1.7)E18	3.12%
CO2	1.1(1.0)E18	2.17%
CH2OH	7.8(4.2)E17	1.55%
CH4	3.3(3.0)E17	0.65%
OH	1.7(0.4)E16	0.03%

*Abundance as described in text for fit in Fig 3C

- Broadband IR spectrometer with HgCdTe and compact line separation (LVF)
- Compact microcrycooler to ≤ 120K to provide long wavelength coverage
- compact optics box designed to remain below 220K
- OSIRIS Rex OVIRS heritage design



Property	Ralph	BIRCHES
Mass kg	11	2
Power W	5	<5#
Size cm	49x40x29*	10x10x15
*19.5x16.0x11.6 inches equivalent		
#includes 3W cryocooler		



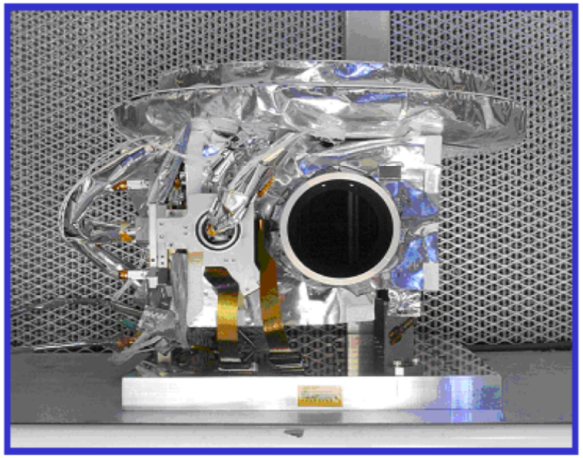
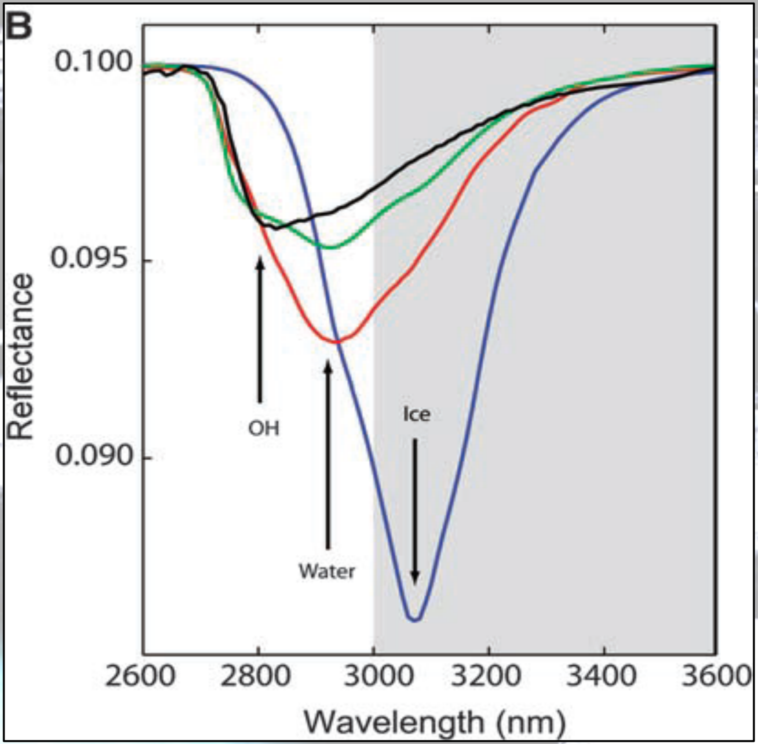


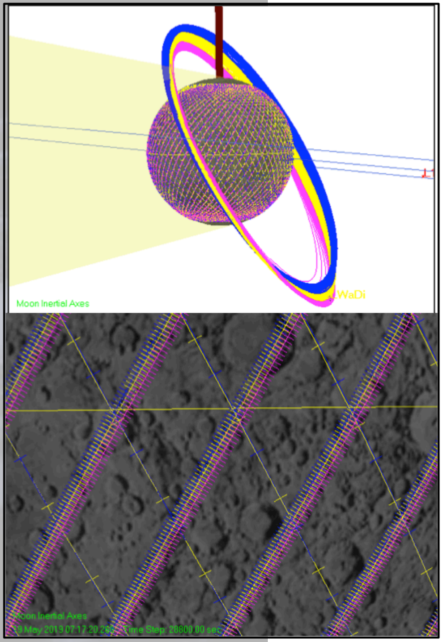
Table C.1 Water-, Volatile-, and Mineral-Related Bands

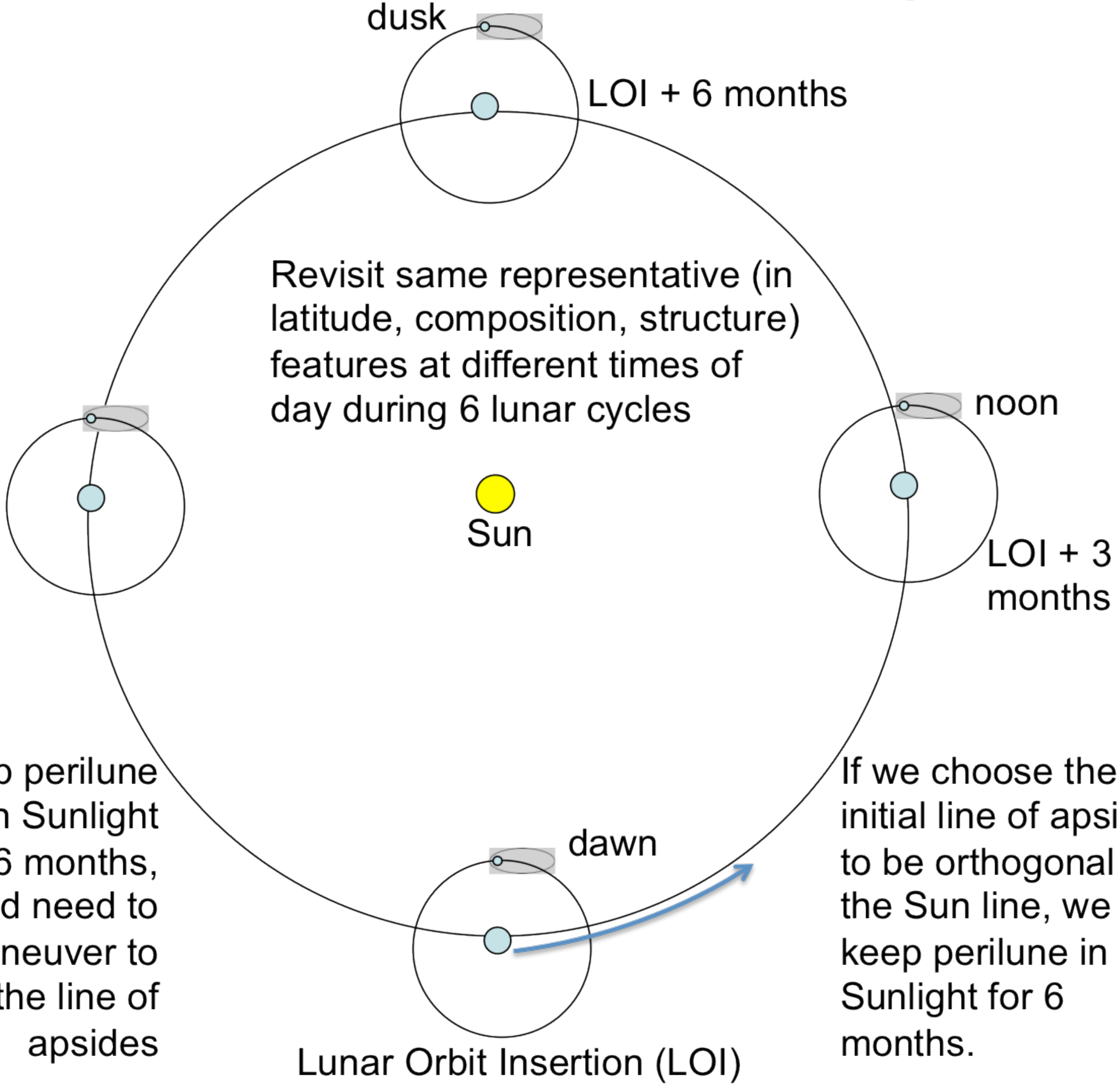
Species	µm	description
Water Form, Component		
water vapor	2.738	OH stretch
	2.663	OH stretch
liquid water	3.106	H-OH fundamental
	2.903	H-OH fundamental
	1.4	OH stretch overtone
	1.9	HOH bend overtone
	2.85	M3 Feature
	2.9	total H2O
hydroxyl ion	2.7-2.8	OH stretch (mineral)
	2.81	OH (surface or structural) stretches
	2.2-2.3	cation-OH bend
	3.6	structural OH
bound H2O	2.85	Houck et al (Mars)
	3	H2O of hydration
	2.95	H2O stretch (Mars)
	3.14	feature w/2.95
adsorbed H2O	2.9-3.0	R. Clark
ice	1.5	band depth-layer correlated
	2	strong feature
	3.06	Pieters et al
Other Volatiles		
NH3	1.65, 2. 2.2	N-H stretch
CO2	2, 2.7	C-O vibration and overtones
H2S	3	
CH4/organics	1.2, 1.7, 2.3, 3.3	C-H stretch fundamental and overtones
Mineral Bands		
pyroxene	0.95-1	crystal field effects, charge transfer
olivine	1, 2, 2.9	crystal field effects
spinels	2	crystal field effects
iron oxides	1	crystal field effects
carbonate	2.35, 2.5	overtone bands
sulfide	3	conduction bands
hydrated silicates	3-3.5	vibrational processes
anticipate wavelength of peak for water absorption		
band would be structural<bound<adsorbed<ice		



Ice Cube measurements will not cut off (Pieters et al. 2009) but encompass the broad 3 um band to distinguish overlapping OH, water, and ice features.

LWaDi 6 Month Mission Concept

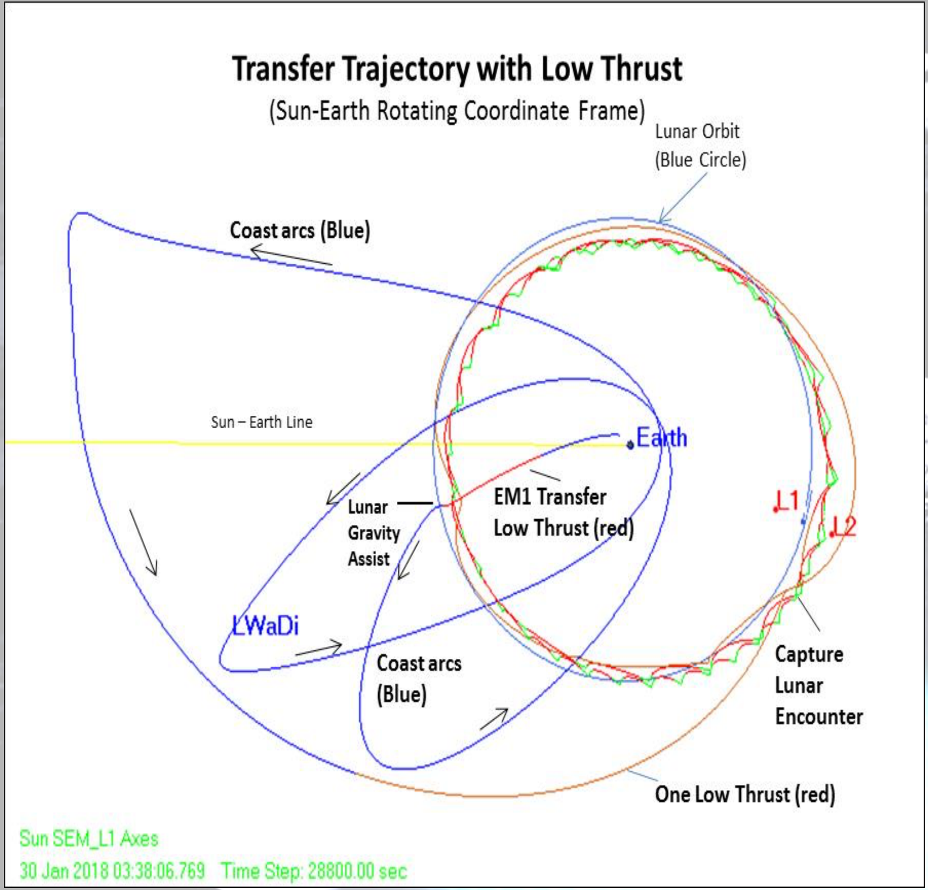




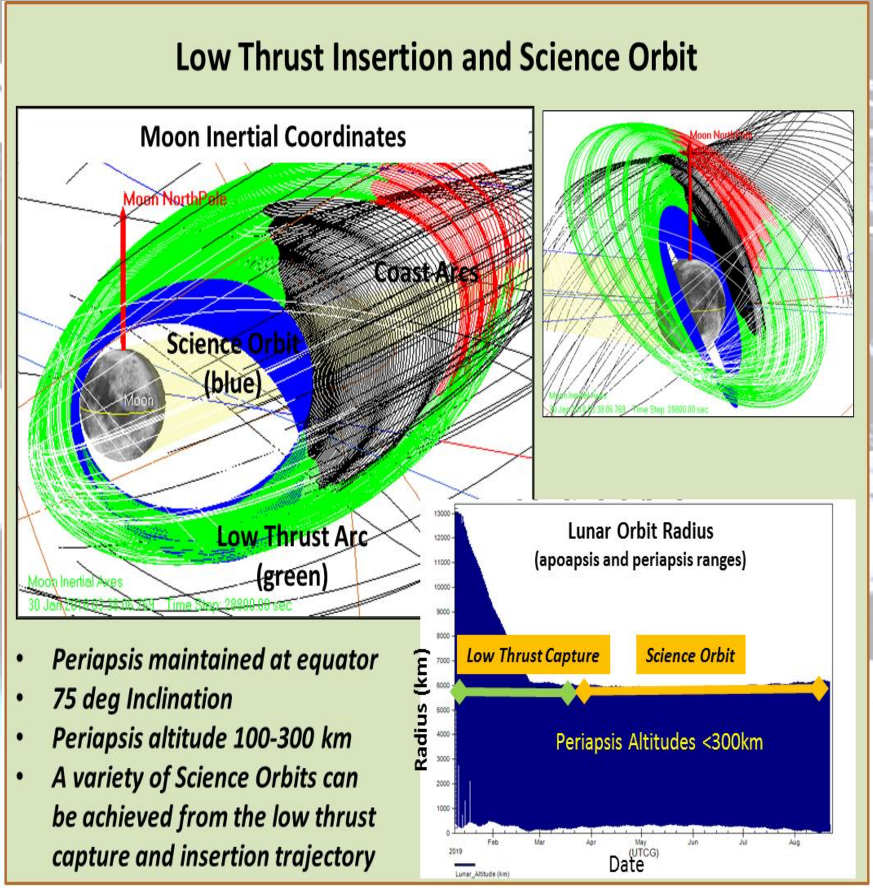
Revisit same representative (in latitude, composition, structure) features at different times of day during 6 lunar cycles

To keep perilune in Sunlight beyond 6 months, we would need to maneuver to rotate the line of apsides

If we choose the initial line of apsides to be orthogonal to the Sun line, we can keep perilune in Sunlight for 6 months.



Low Thrust Insertion and Science Orbit



- *Periapsis maintained at equator*
- *75 deg Inclination*
- *Periapsis altitude 100-300 km*
- *A variety of Science Orbits can be achieved from the low thrust capture and insertion trajectory*

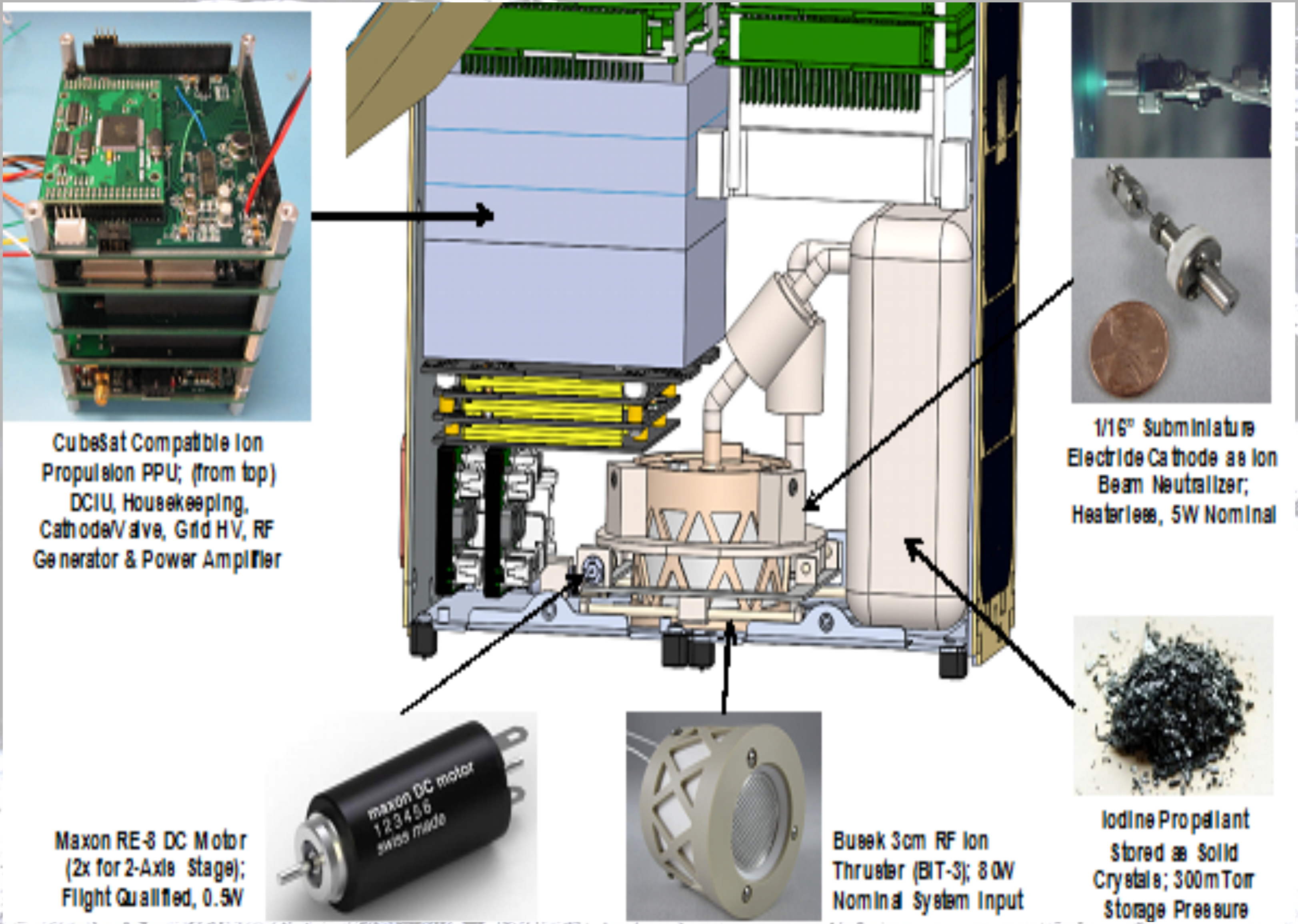
IceCube utilizes a minimal DV transfer trajectory harnessing expertise of GSFC flight dynamics.

IceCube lunar capture and science orbit designed by experienced GSFC flight dynamics team.

IceCube versus Previous Missions

Mission	Finding	IceCube
Cassini VIMS, Deep Impact	surface water detection, variable hydration	water & other volatiles, fully characterize 3 µm region as function of several times of day for same swaths over range of latitudes w/ context of regolith mineralogy and maturity, radiation and particle exposure, for correlation w/ previous data
Chandra M3	H2O and OH (<3 microns) in mineralogical context nearside snapshot at one lunation	
LCROSS	ice, other volatile presence and profile from impact in polar crater	
LP, LRO, LEND LAMP DVNR LOLA LROC, LADEE	H+ in first meter (LP, LEND) & at surface (LAMP) inferred as ice abundance via correlation with temperature (DIVINER), PSR and PFS (LROC, LOLA), H exosphere (LADEE)	

Busek Iodine ion propulsion system



CubeSat Compatible Ion Propulsion PPU; (from top) DCIU, Housekeeping, Cathode/Valve, Grid HV, RF Generator & Power Amplifier

1/16" Subminiature Electride Cathode as Ion Beam Neutralizer; Heaterless, 5W Nominal

Maxon RE-8 DC Motor (2x for 2-Axis Stage); Flight Qualified, 0.5W

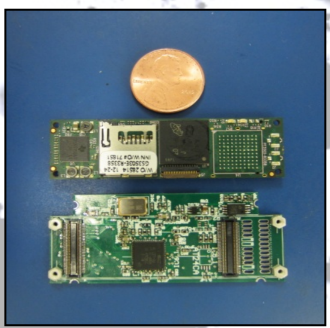
Busek 3cm RF Ion Thruster (BIT-3); 80W Nominal System Input

Iodine Propellant Stored as Solid Crystals; 300mTorr Storage Pressure

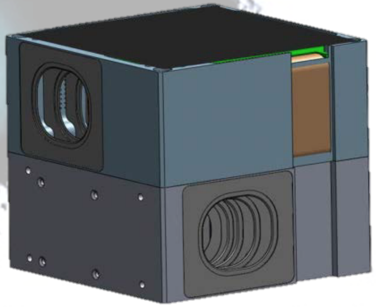
Bus Components

Thermal Design: with minimal radiator for interior the small form factor meant that interior experienced temperatures well within 0 to 40 degrees centrigrade, except for optics box which has a separate radiator.

Communication, Tracking: X-band, JPL Iris Radio, dual X-band patch antennas, X-band dish (trade availability, cost, dB, and DSN compatibility, live with the fact this hasn't flown in deep space)



C&DH: very compact and capable Honeywell DM microprocessor, at least one backup C&DH computer (trade volume, complexity, cubesat heritage, live with the fact this hasn't flown in deep space)



GNC/ACS: multi-component (star trackers, IMU, RWA) packages with heritage available, including BCT XB1, which can interface with thrusters (trade cost, volume, cubesat heritage, live with the fact this hasn't flown in deep space)

IceCube Concept: Morehead CubeSat Bus

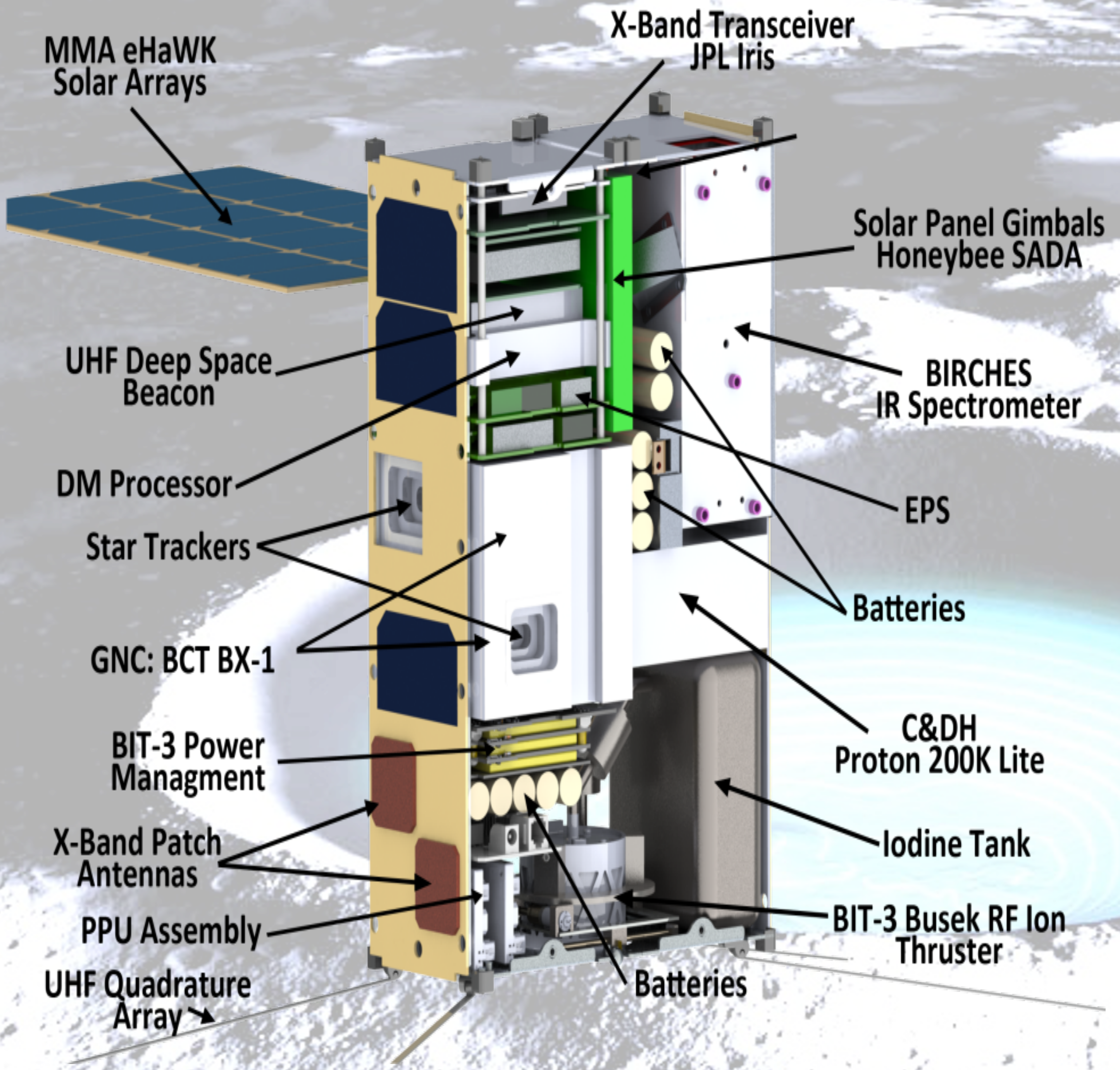


Table 4: Lunar IceCube Subsystems

System	Mass	Volume (in Us)	Power Use	Rad Tolerance	Dollars*	Source
Structures, Thermal Management	1.2 kg	6U Exterior & Rings	N/A	N/A	69K	MSU
C&DH: Proton 200K Lite/ Custom Daughter BCT XB1 for ACS Control	0.36 kg	0.75 U	5 W	>100 krad	240K	Space Micro BCT XB1 MSU
Harnesses, cables, coatings, elastomerics	0.5 kg	Conformal w/in structure	N/A	N/A	12K	MSU
Power (Solar panels & Gimbals MMA HaWK 72 W Array	0.340kg x 2 Deployed 0.190kg x 2 Fixed (Side)	Deployable panels intrude 10 mm into structure (each)	N/A	TBD	185K	MSU + MMA
Solar Panel Drive Articulators HoneyBee SADA	0.40 kg x 2	10 x 10 x 0.65 cm (Each) (0.25U)	5 W	10 krad	Included Above	Honeybee Robotics-MMA
EPS + Batteries MSU + TBD	2.4 kg	0.5 U	Quiescent Draw = 10 mW	> 10 krad	36K	MSU
Propulsion: Busek BIT-3cm RF Ion Iodine	2.5 kg	2U	60W	TBD	1,000K	Busek
ACS/GNC: BCT XB1	2.1 kg	0.75 U	6.3W Cont.	TBD	250K	BCT XB1
Comms: JPL Iris	0.5 kg	0.5U + Antennas	12.8 W-Transponder 6.4 W- Receive	50 krad	500 K	JPL IRIS
IR Spectrometer	0.62 kg	1.5 U	<5 W	TBD	In budget	GSFC
Payload Processor: DM	0.350 kg	10 x 10 x 4 cm (0.25U)	2 W per processor continuous	Multiple processors (8) & middleware	68K	MSU/Honey well

The Next Frontier: CubeSats for Deep Space

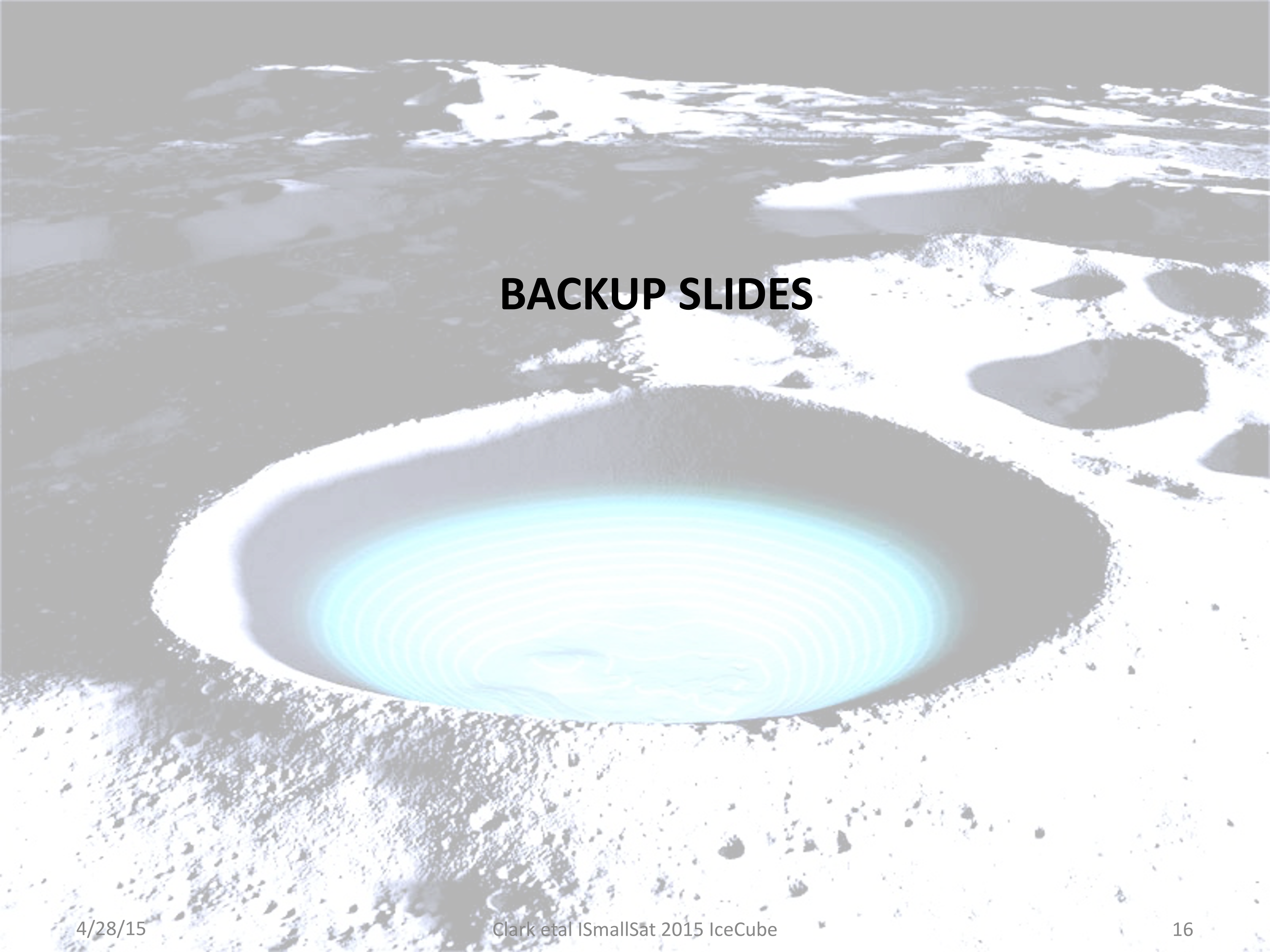
3rd International Workshop on LunarCubes

November 13-15, 2013 – Palo Alto, CA

Lunar Science Illuminating the Universe

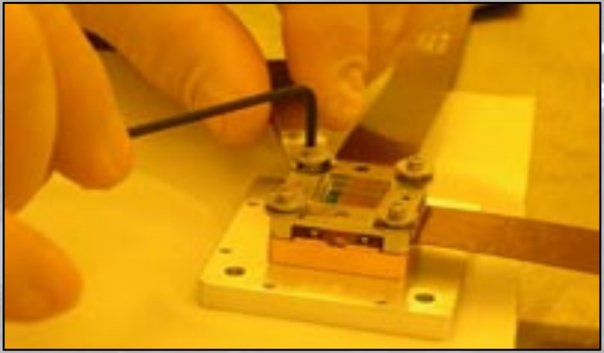
1st International Workshop on Scientific Opportunities in Cislunar Space

November 9th, 2014 - Tucson, AZ

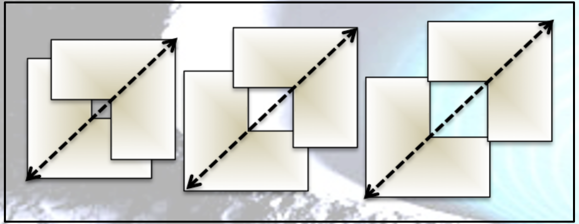


BACKUP SLIDES

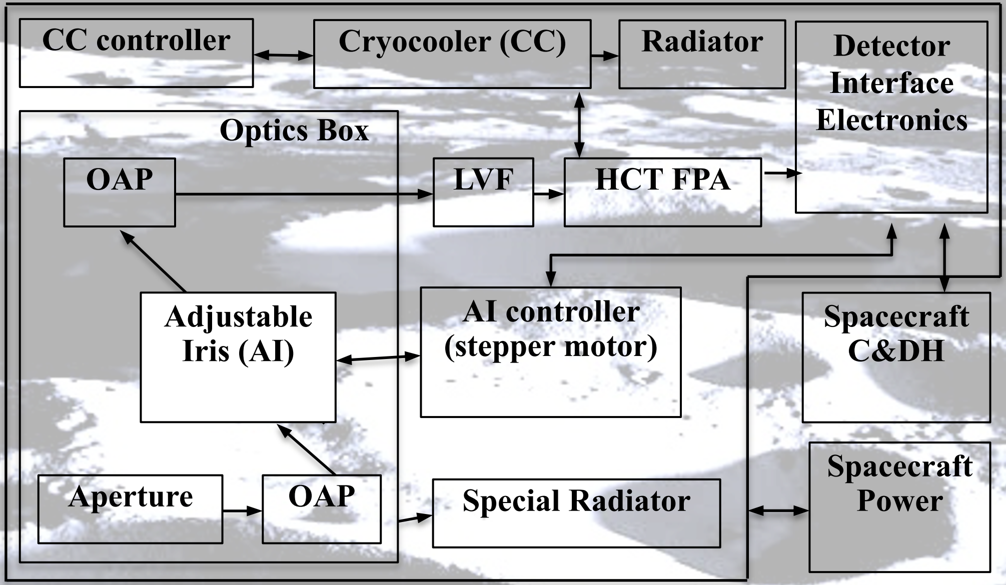
Spectrometer Components



BIRCHES utilizes a compact Teledyne H1RG HgCdTe FPA and JDSU linear variable filter detector assembly leveraging OSIRIS REx OVIRS.



Adjustable Iris maintains footprint size at 10 km by varying FOV regardless of altitude



CC controller

Cryocooler (CC)

Radiator

Detector Interface Electronics

Optics Box

OAP

LVF

HCT FPA

Adjustable Iris (AI)

AI controller (stepper motor)

Spacecraft C&DH

Aperture

OAP

Special Radiator

Spacecraft Power

BIRCHES block diagram illustrates simplicity and flexibility of design.

Off the shelf tactical cryocooler with cold finger to maintain detector at ≤140K

Bus Components

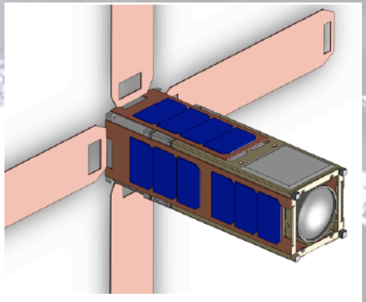
Power:

compactly packaged Li-based batteries (e.g., GOM) that provide adequate power storage for longest 'eclipse' of sun in orbit;

electrical power system, for which many cubesat heritage options are available

Deployable solar panels, for which a number of choices are available (from top to bottom, turkey tail, cross, table, gimballed version of cross). Producers include MMA Design, Honeybee Robotics) Require >50 W running an active propulsion system, which should be more than adequate for other needs when propulsion system isn't running.

trade space cost, mass, reliablity, although volume of solar panels is in the 'cheat space' and doesn't count against 6U total



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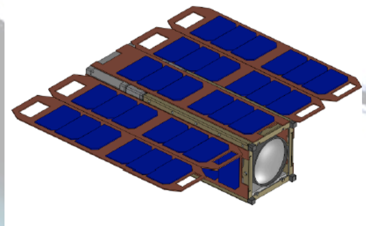

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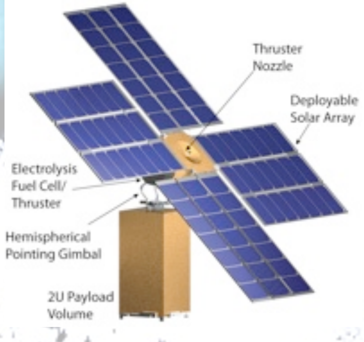

EM1 Selectees to Date

Program	Target, Description	Payload	Lead
HEOMD NEXT STEP	Moon, orbiter, Ice Cube	broadband IR cryocooled.	Morehead State U/NASA GSFC/Busek
HEOMD AES	Lunar Flashlight orbiter. (Surface ices in permanently shadowed 'cold traps')	NIR instrument.	JPL
HEOMD AES	Near Earth Asteroid Scout	Imager to characterize asteroid dynamics and surface	JPL
HEOMD AES	BioSentinel	Radiation Exposure Induced Genetic Damage Experiment	NASA/Ames
HEOMD NEXT STEP	??	???	Lockheed Martin
SMD SIMPLEx			
STMD Centennial Challenges			

Overarching Question: Considering the science priorities and resulting range of science investigations, and the range of potential payloads, what does a 'lunarcube' platform look like? 6U, needs robust propulsion system (>1.5 km/sec delta V) mostly to achieve desired orbit from lunar capture, can carry up to 2U payload, >60W power desirable, needs robust thermal protection design, requires 1 year plus

Design Challenge 2 Cubesat Concepts to Cubesat Missions: Applied to this concept

1) Overview science, investigation, operational concept and principal drivers (needs) volatiles study, low periapsis elliptical inertially orbit, thermal design and mobility

2**) Trade space (prioritized needs for which optimized capabilities are needed versus resources (volume, cost, bandwidth) available)?** More robust and compact ACS, Comm, Power systems available but at cost.

3) What are your perceived performance limitations, risks, and descope options? Bandwidth limited (comm), delta V limited (but focused mission achievable), thermal design challenging and would be improved with ‘smart’ materials and mechanisms. Radiation exposure risk involving use of RHBD hardware/software and more reliable parts sources. Descope involves taking data on way down to final orbit, baseline is 3 months (rather than 6) in that orbit.

4) Link science objective, measurement, instrument requirements, mission requirements, science product (see chart)